

# Neutrino Physics

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and

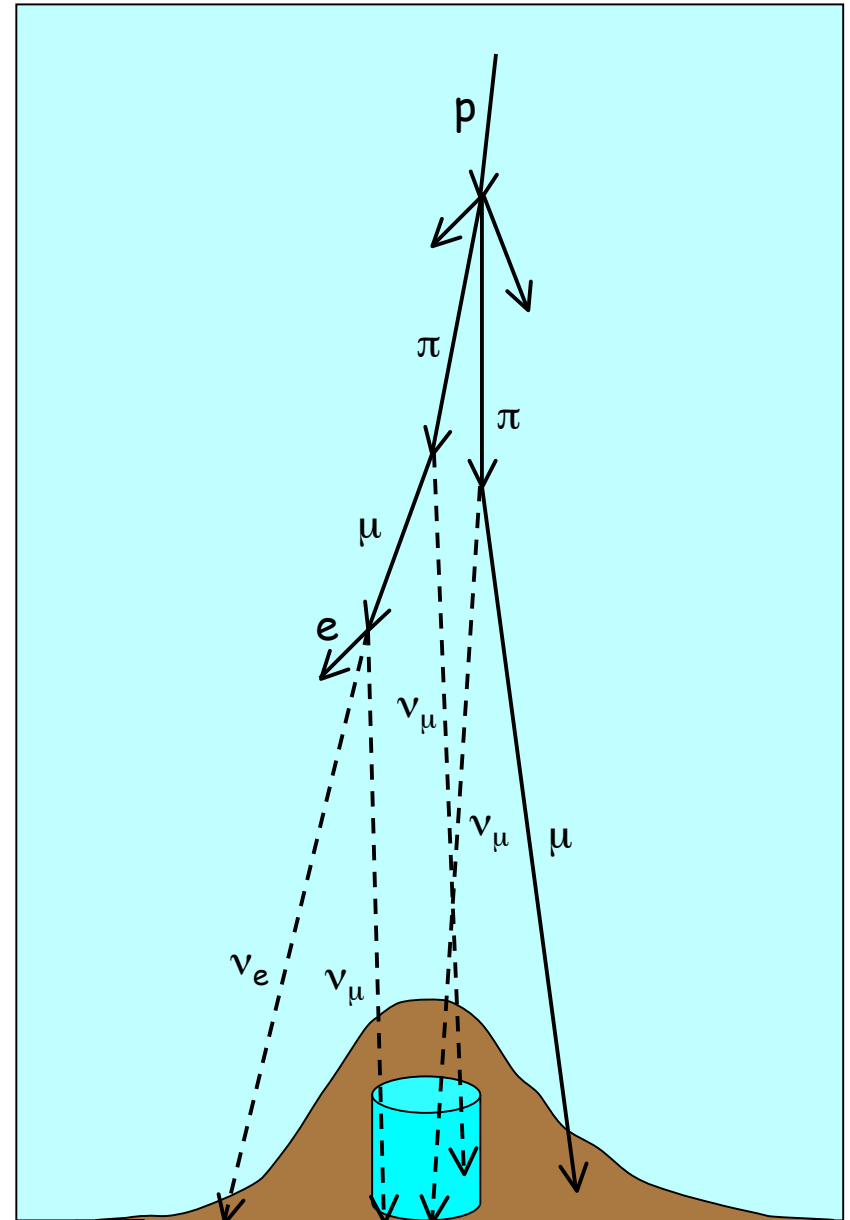
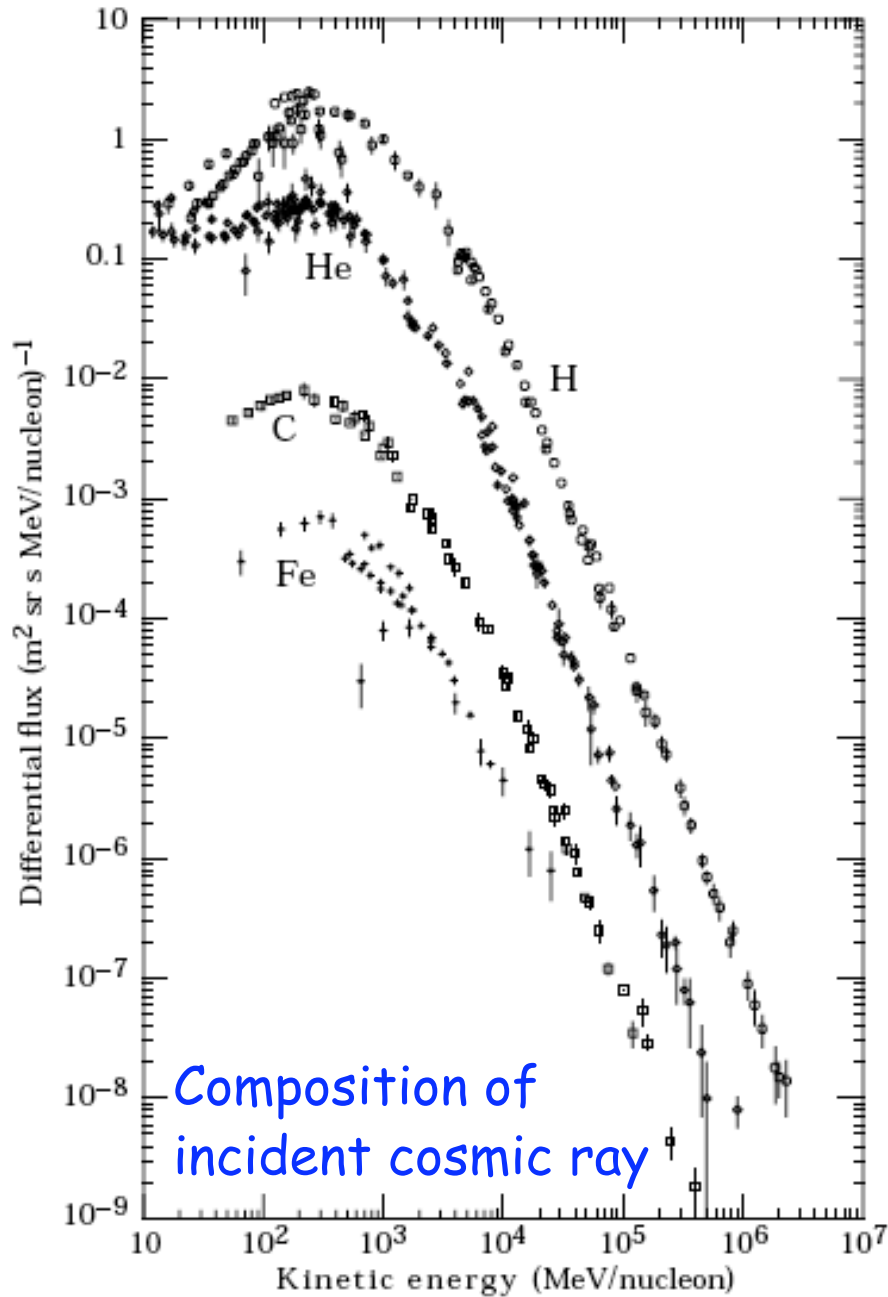
Lawrence Berkeley National Laboratory

Lecture 9, 14 June, 2007

# Outline

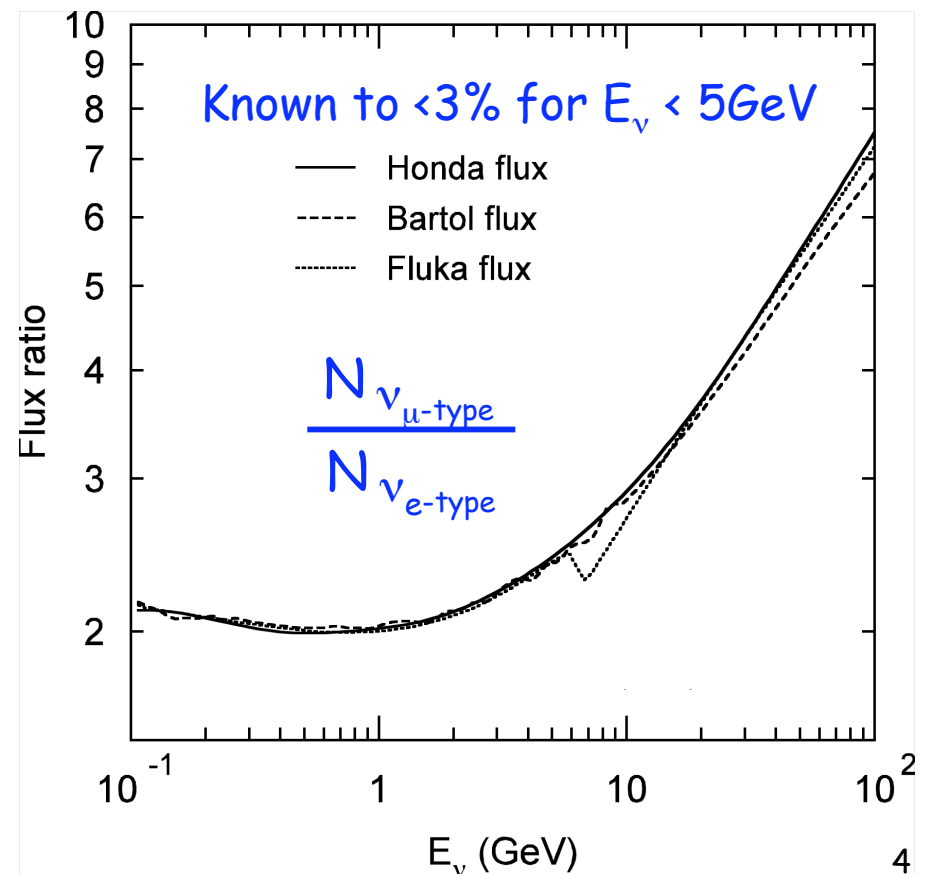
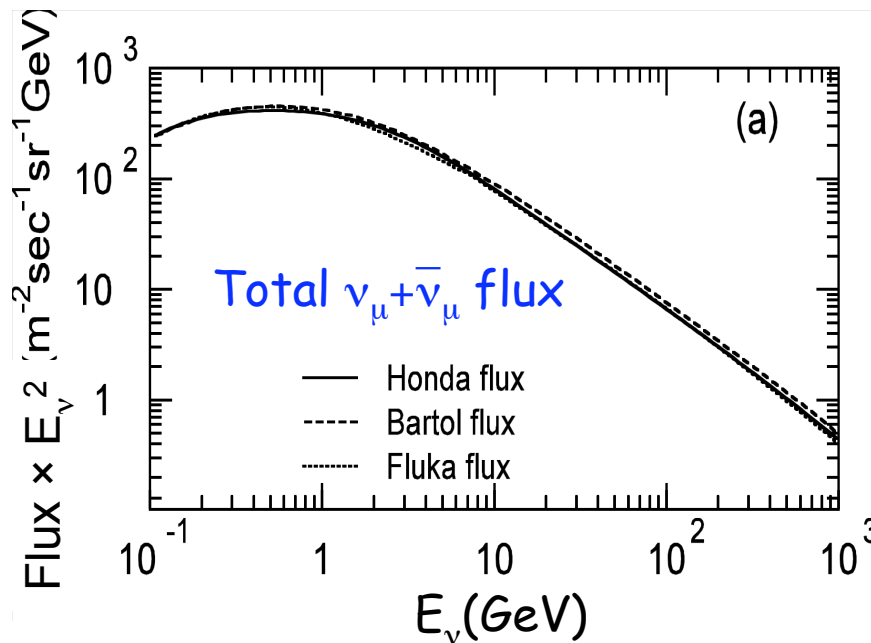
- Study of atmospheric neutrino
  - Super-K
  - Soudan 2
- Confirming atmospheric neutrino oscillation
  - K2K
  - MINOS

# Production Of Cosmic-ray Neutrinos

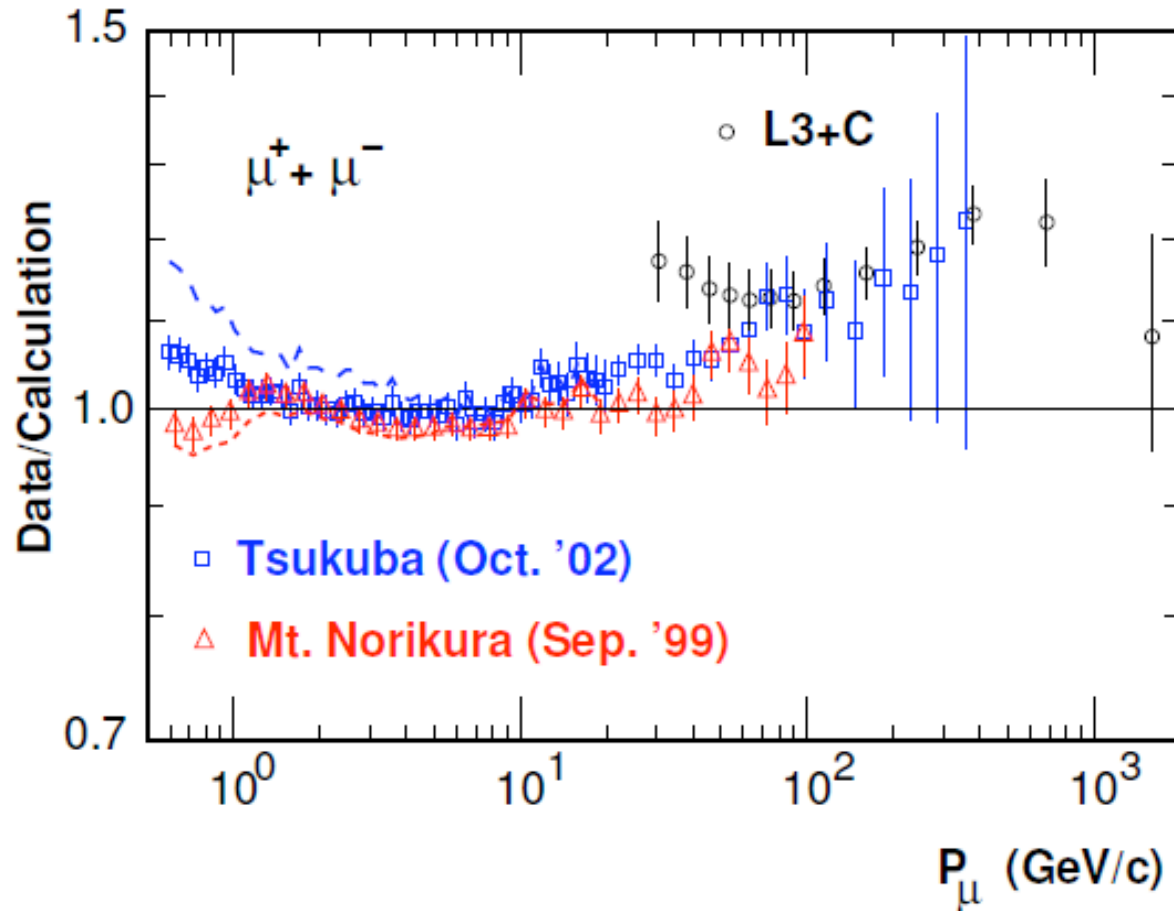


# Estimating Atmospheric Neutrino Flux

- Use measured cosmic-ray intensity at upper atmosphere and cross section of p-nucleon interaction to calculate the numbers of  $\pi$  and K produced.
- Transport the mesons through the atmosphere, taking the effect of the geomagnetic field into account and allow them to decay.



# Cross Check Calculation With Cosmic-ray Muons

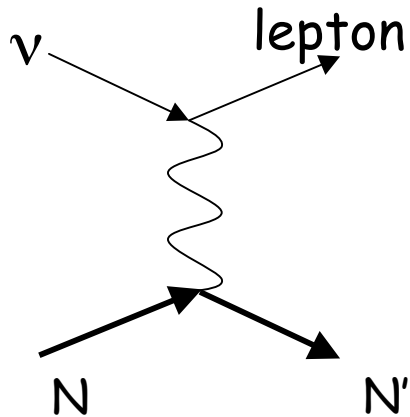


Calculation is good to better than 10% for momentum less than  $\sim 100$  GeV

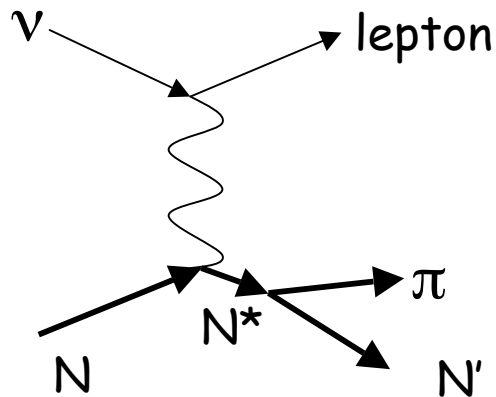
# Neutrino Interactions

More on the charge-current processes:

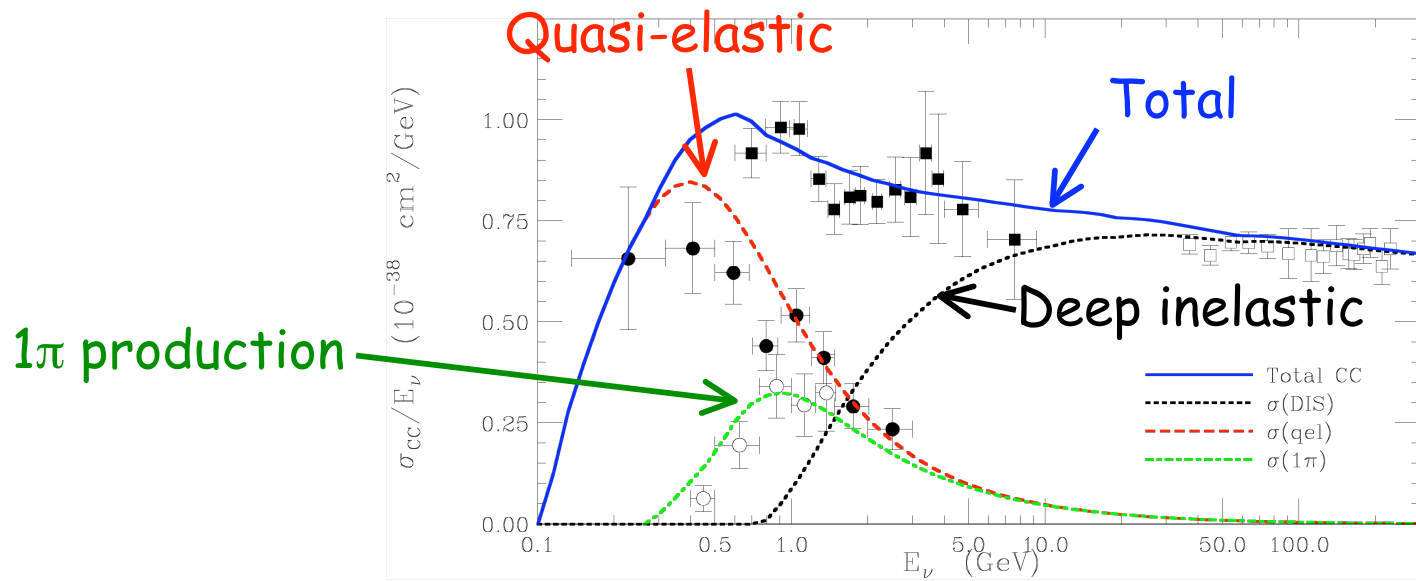
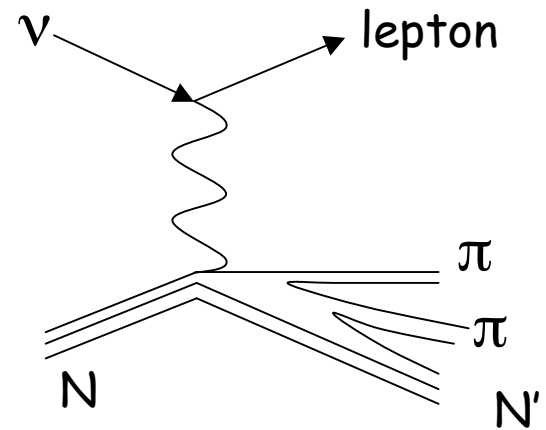
Quasi-elastic



1π production

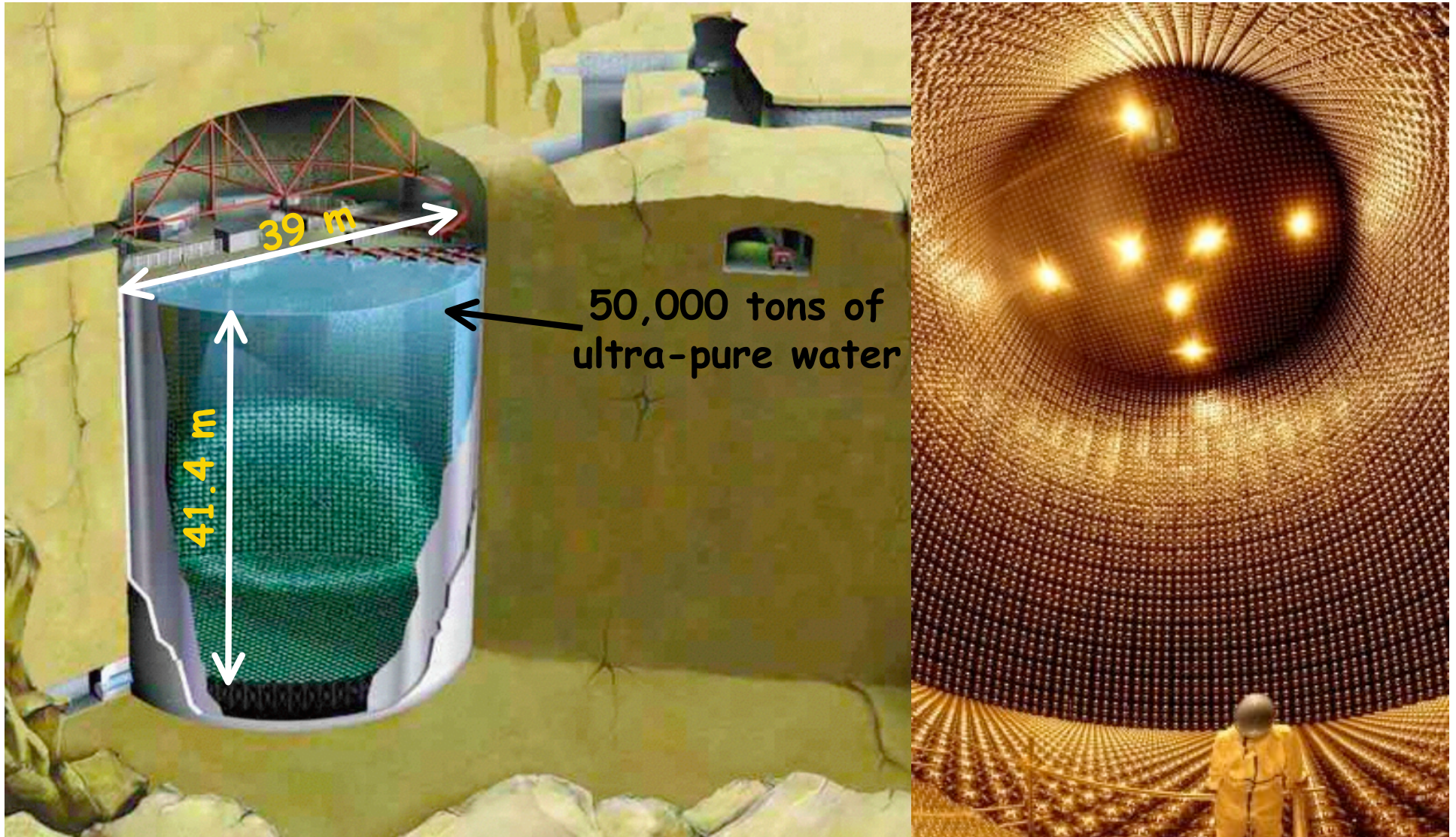


Deep inelastic



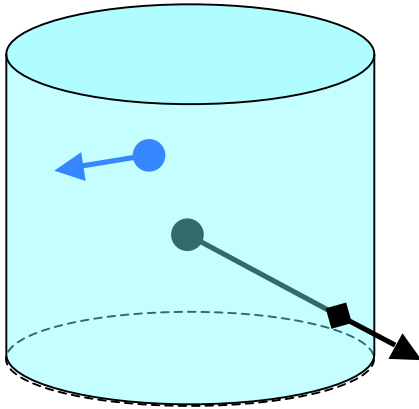
# Study Atmospheric Neutrino With Water Cherenkov

- Kamiokande, and Super-Kamiokande in Japan

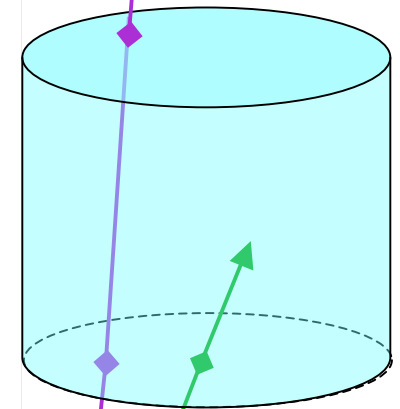
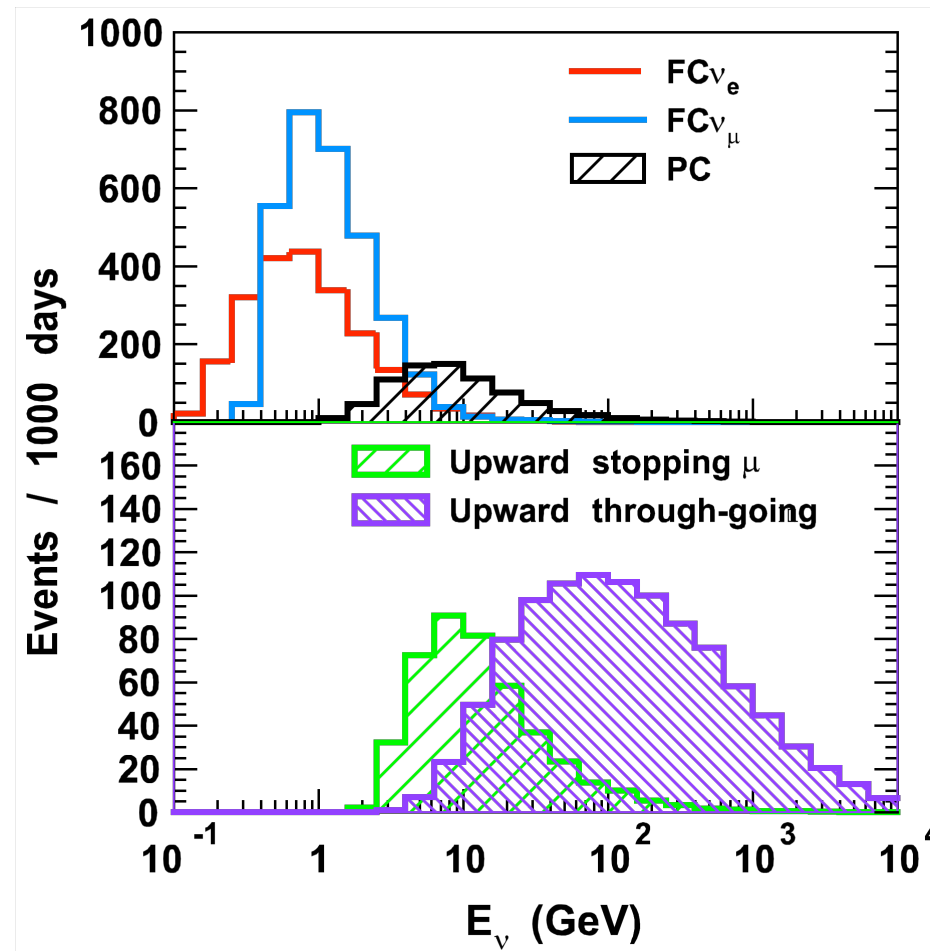


# Classification Of Detected Neutrino Events

Fully Contained (FC)



Partially Contained (PC)



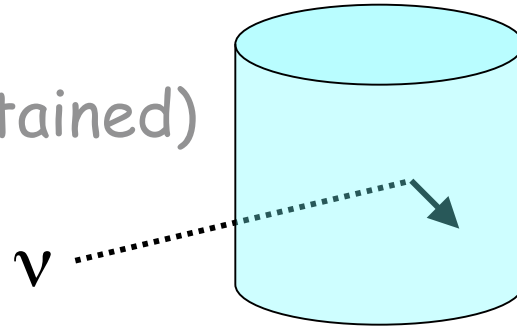
Stopping  $\mu$

Through-going  $\mu$



# Examples Of Charged-current Events

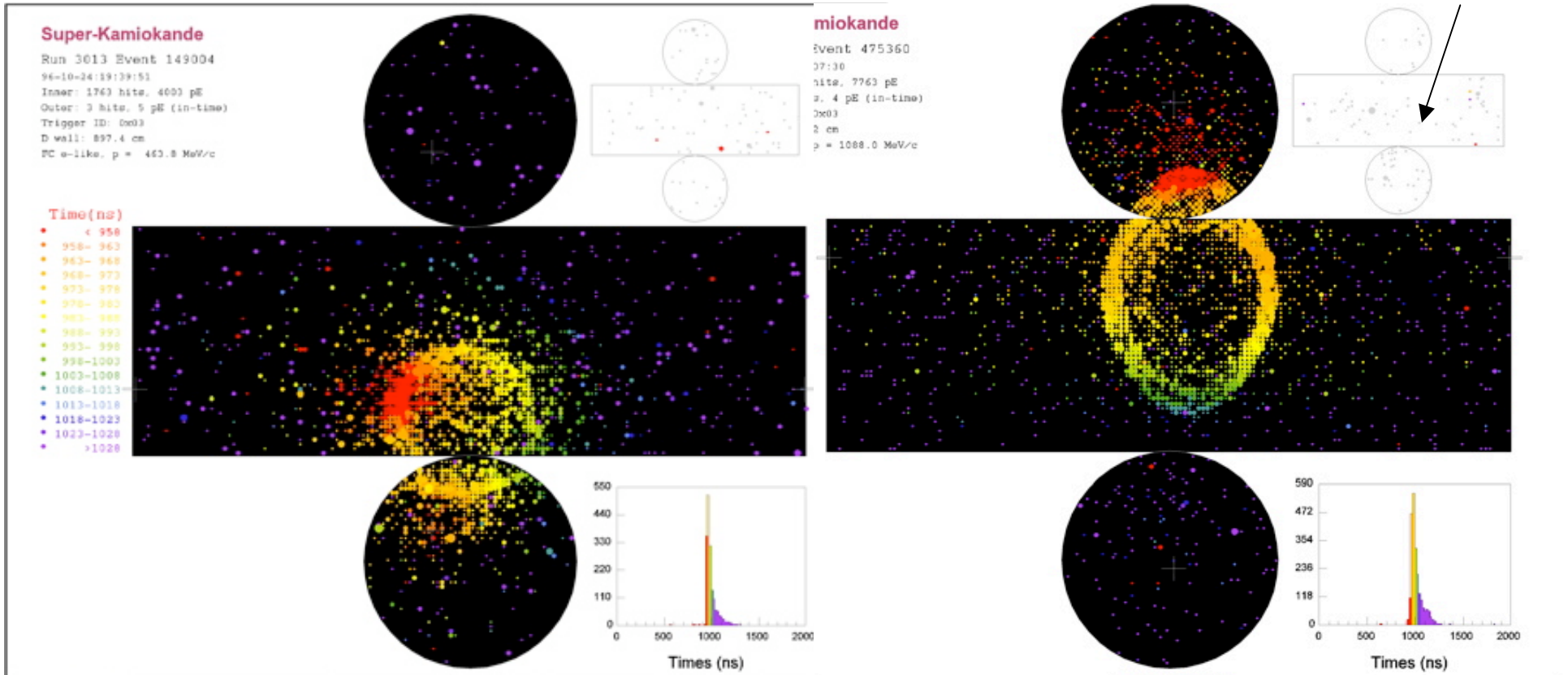
FC  
(fully contained)



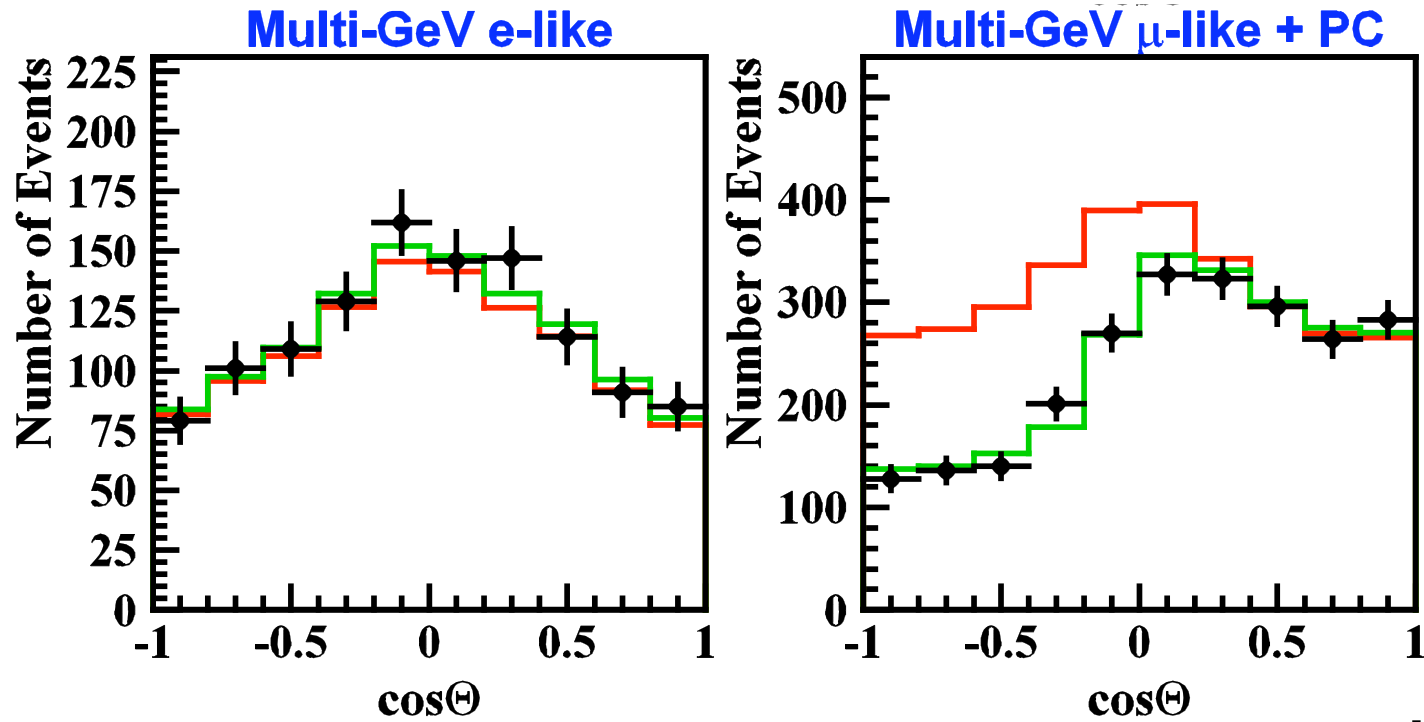
electron-like Cherenkov ring

muon-like Cherenkov ring

Outer detector  
(no signal)



# Surprising Behavior Of $\nu_\mu$

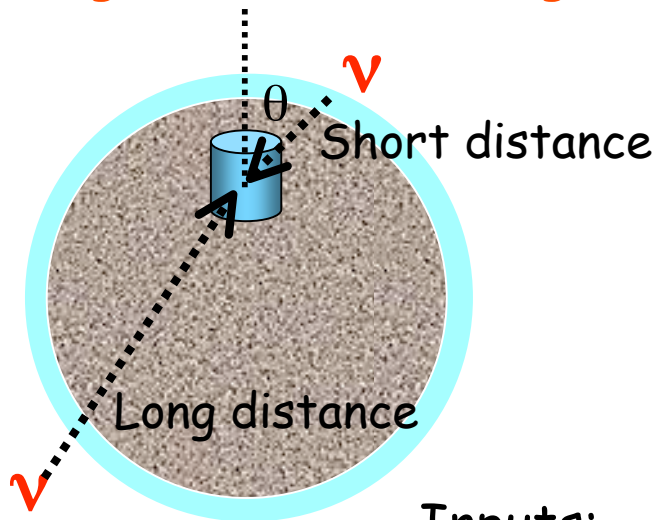


# Interpretation Of Atmospheric $\nu$ Results

- The unexpected results are evidence of neutrino oscillation.
- There is no compelling evidence of  $\nu_\mu \rightarrow \nu_e$  oscillation.
- The disappearance of  $\nu_\mu$  is due to  $\nu_\mu \rightarrow \nu_\tau$  oscillation.
- Adopt 2-flavour model to analyze the data:

$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - \sin^2 2\theta \sin^2 \left( \frac{\Delta m^2 L}{4E} \right)$$

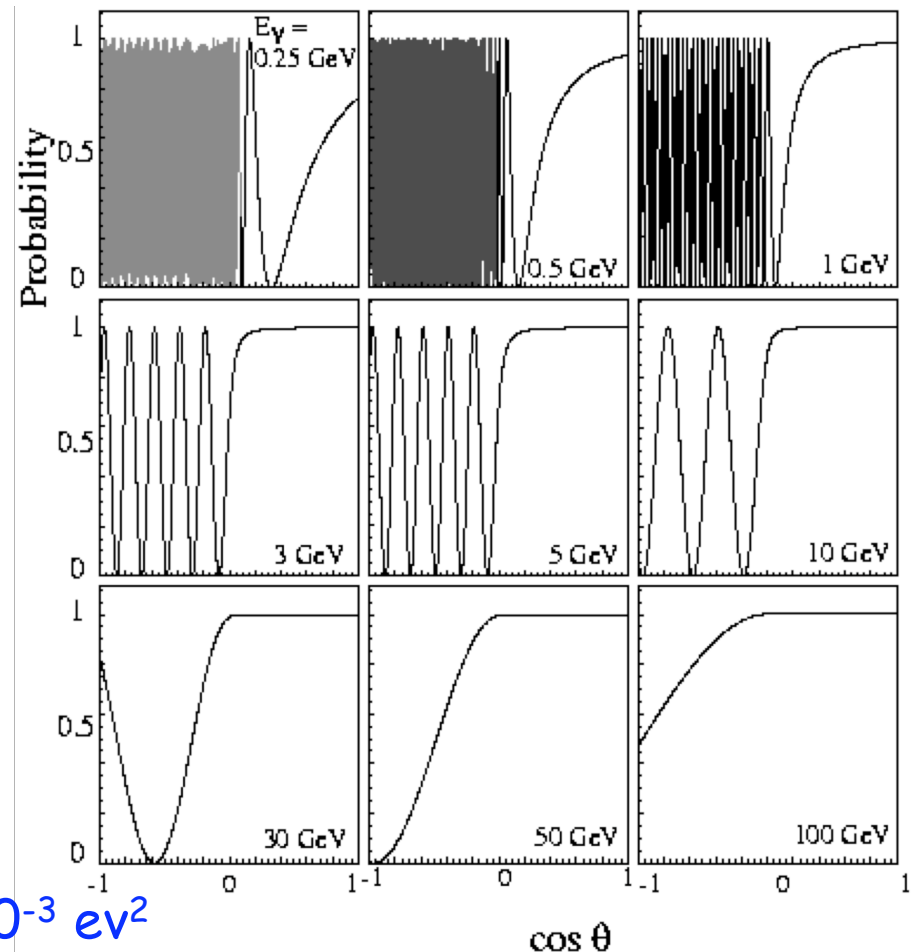
$L$  changes with zenith angle



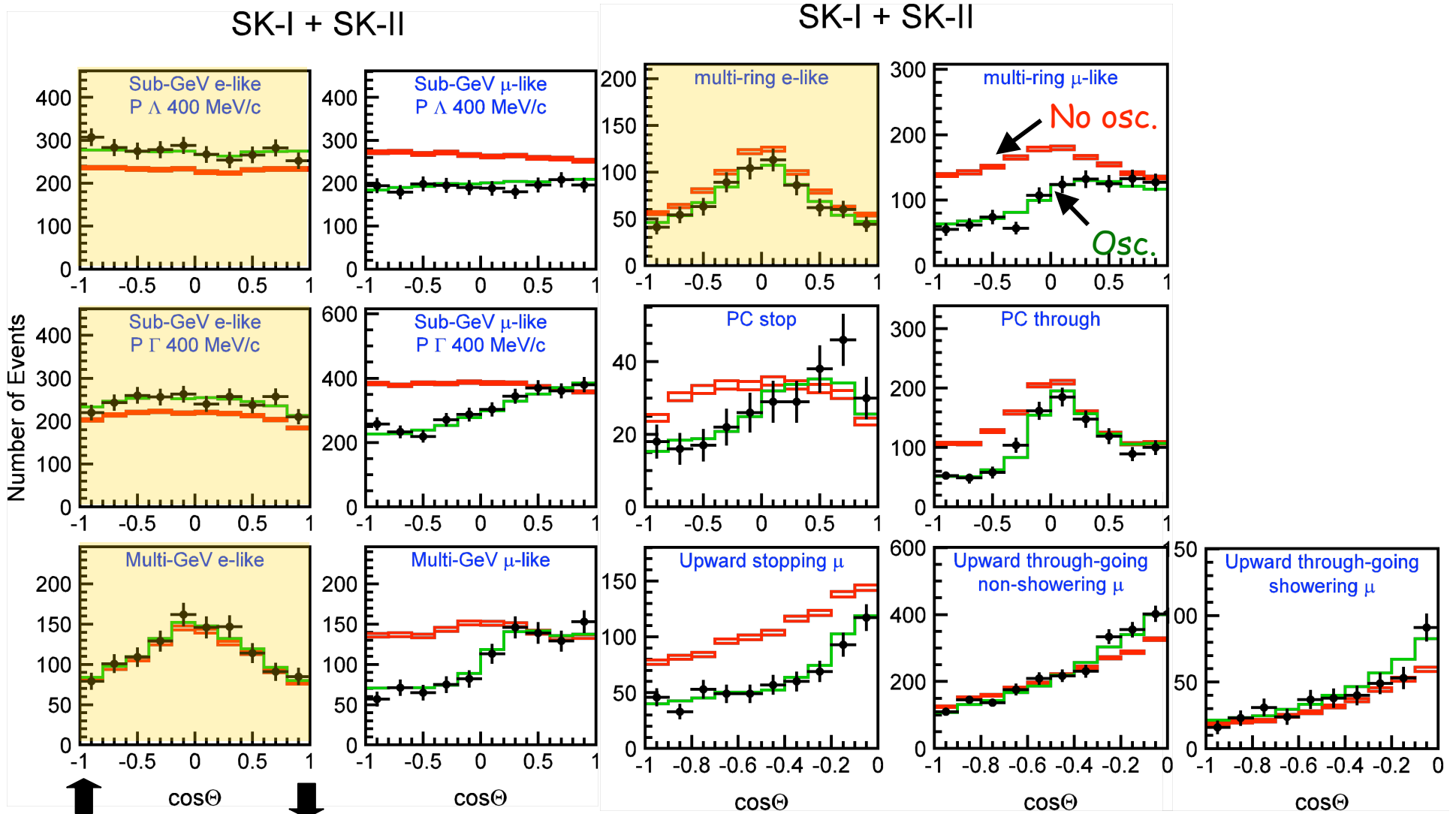
Inputs:

$$\sin^2 2\theta = 1.0,$$

$$\Delta m^2 = 2.5 \times 10^{-3} \text{ eV}^2$$

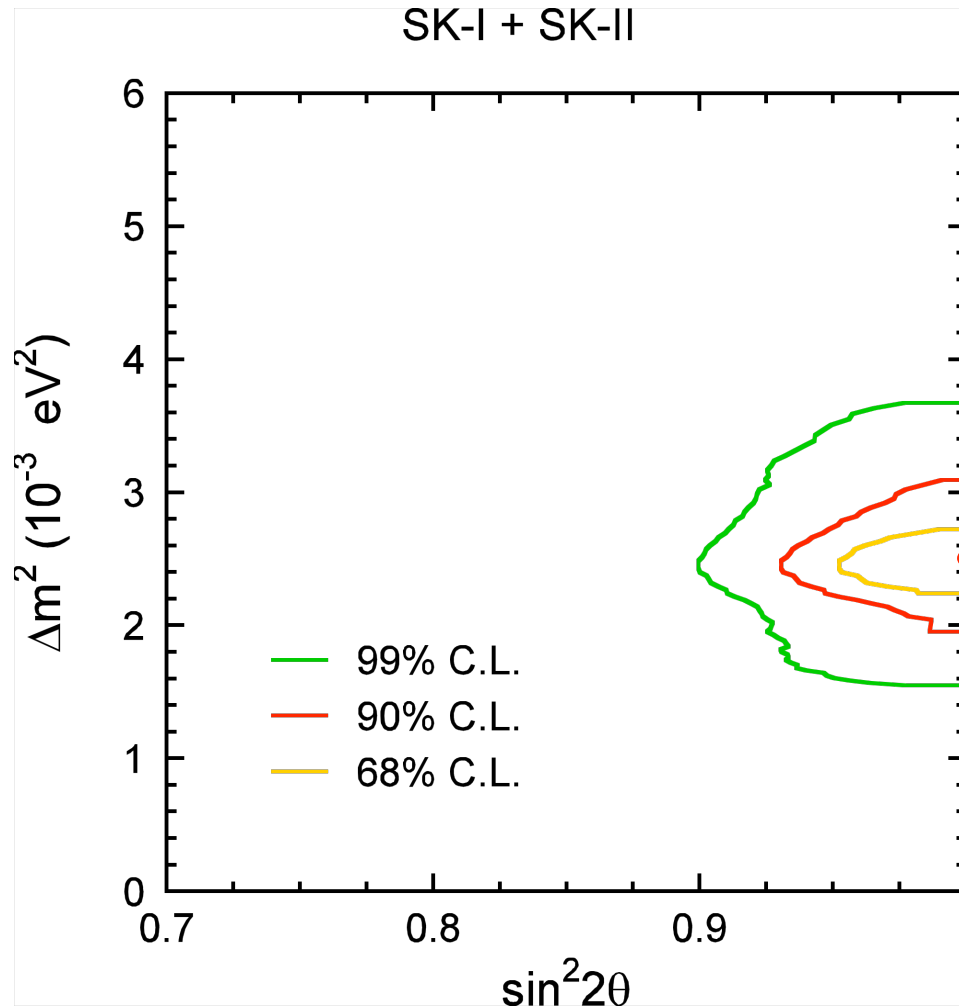


# More Results From Super-K



# Results On Oscillation Parameters

Using number of events only:



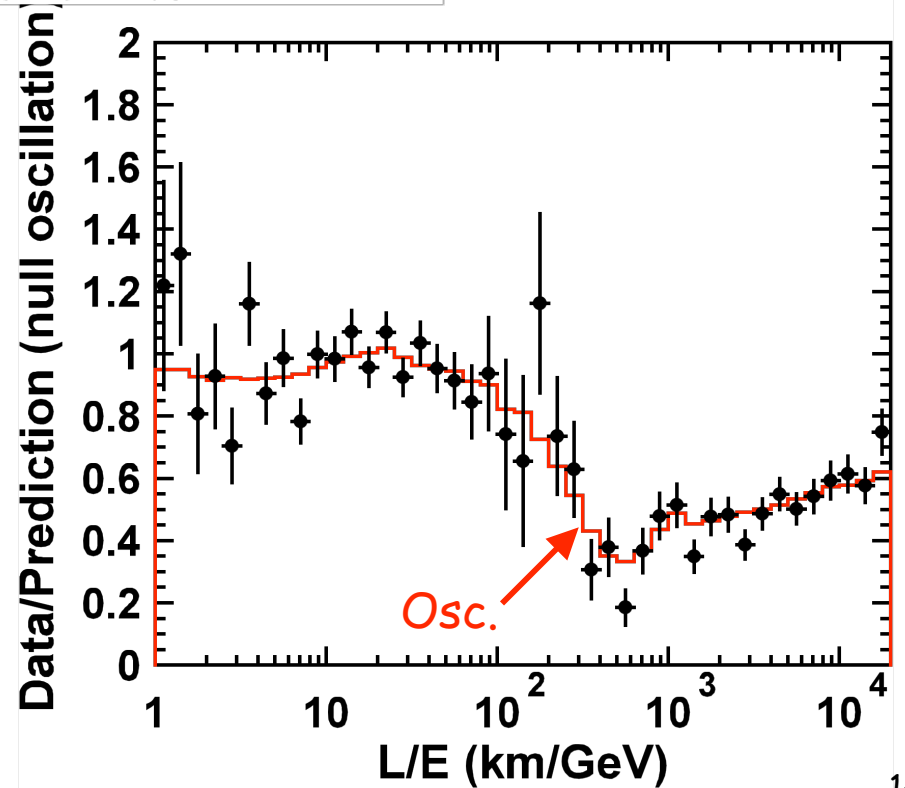
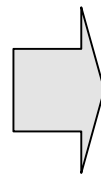
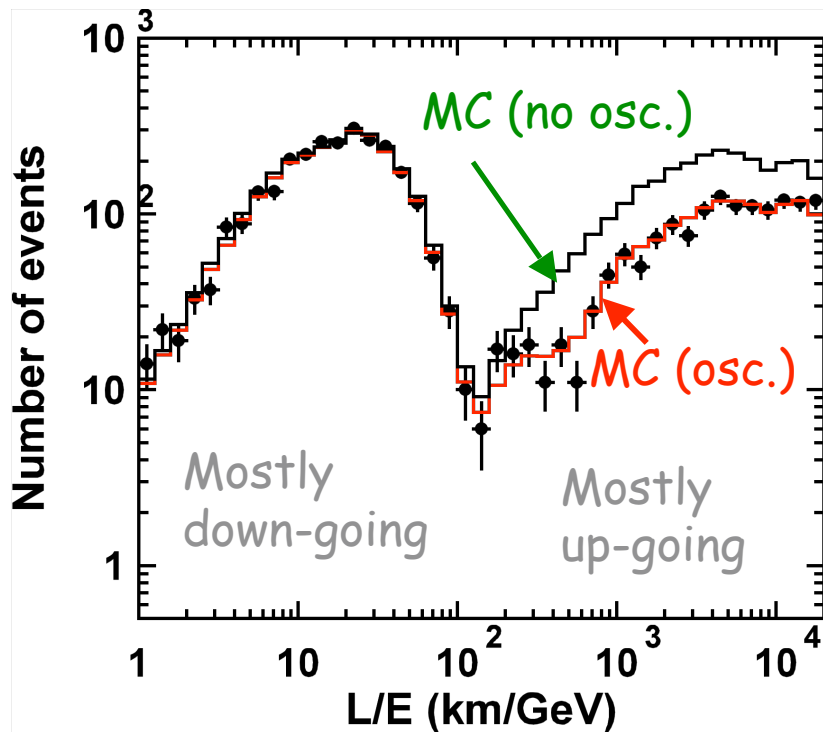
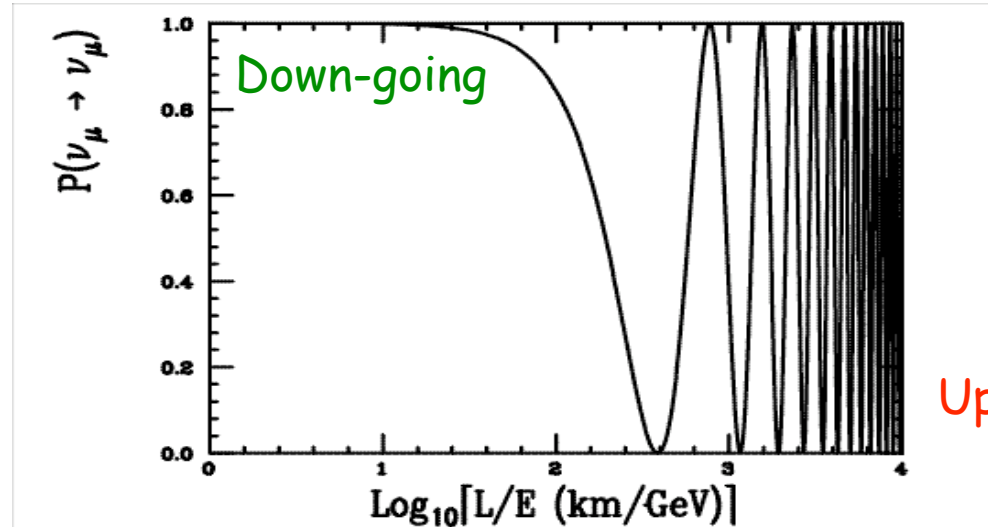
Best Fit:

$$\Delta m^2 = 2.5 \times 10^{-3} \text{ eV}^2$$

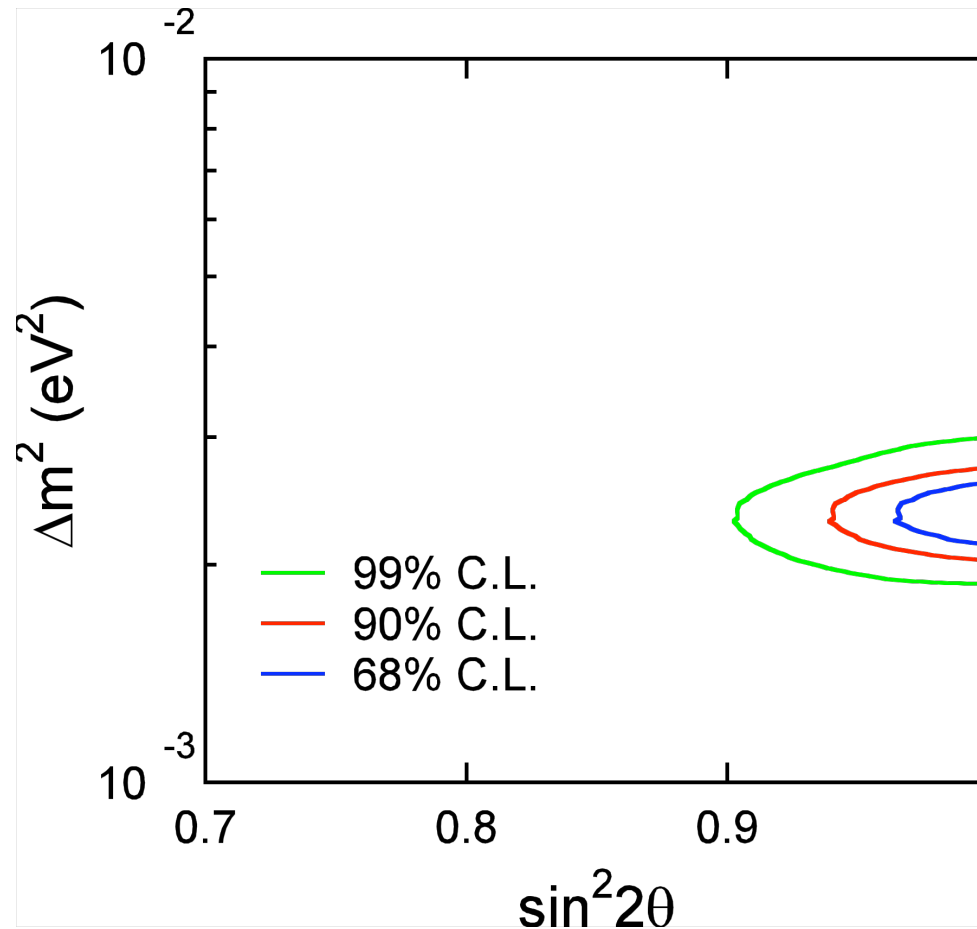
$$\sin^2 2\theta = 1.00$$

$$\chi^2 = 839.7 / 755 \text{ dof}$$

# How About The Oscillation Pattern?



# Results of Atmospheric Neutrino Oscillation

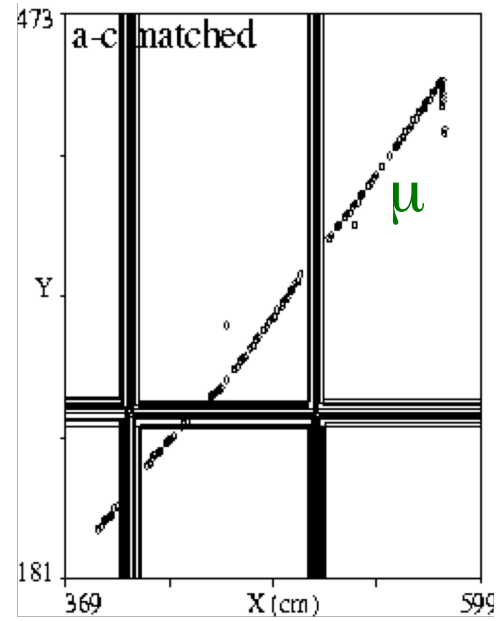


At 90% CL:

$$2.0 \times 10^{-3} \text{ eV}^2 < \Delta m^2 < 2.8 \times 10^{-3} \text{ eV}^2$$

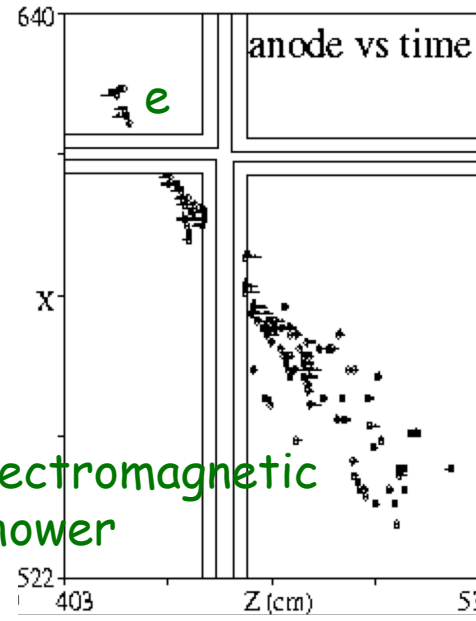
$$\sin^2 2\theta > 0.93$$

# Soudan 2 In U.S.A.

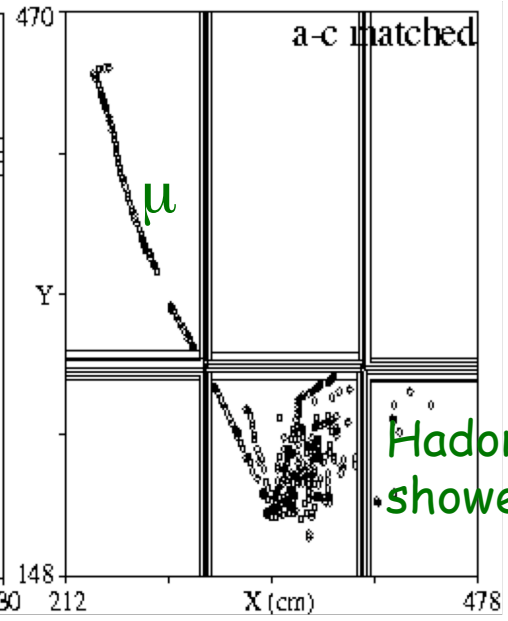


$\nu_\mu$  CC  
Quasi-elastic

$\nu_e$  CC



electromagnetic shower



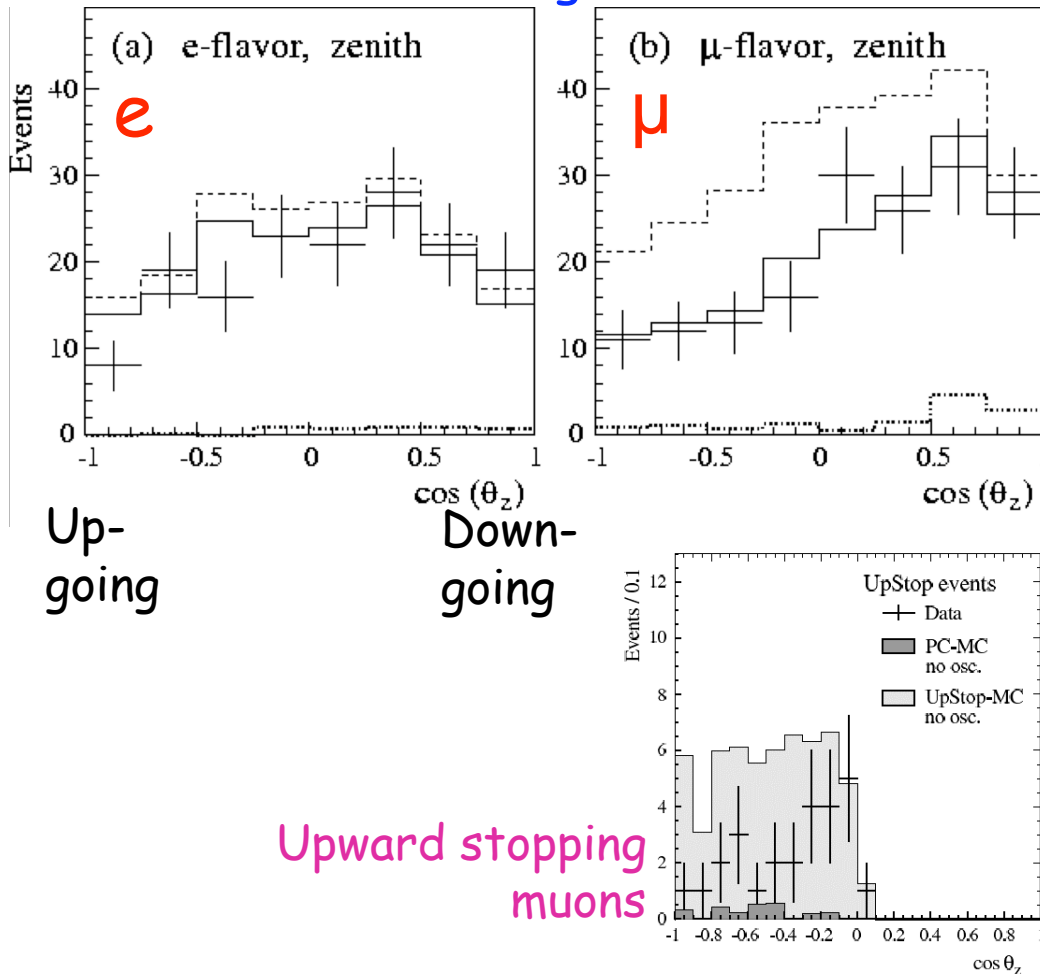
$\nu_\mu$  CC  
Deep inelastic



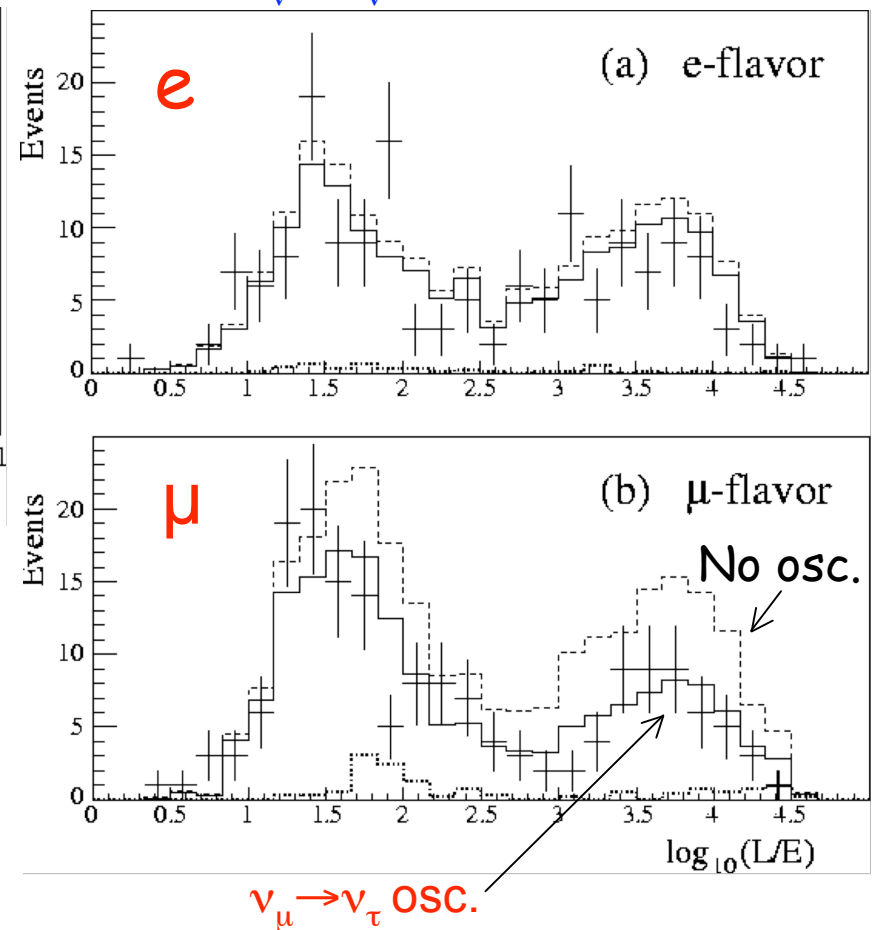
# Confirmation From Soudan 2

- 5.9 kton-yr exposure
- Studied partially-contained and upward stopping muon events.
- Measured  $L/E$  with the high resolution sample

Zenith angle

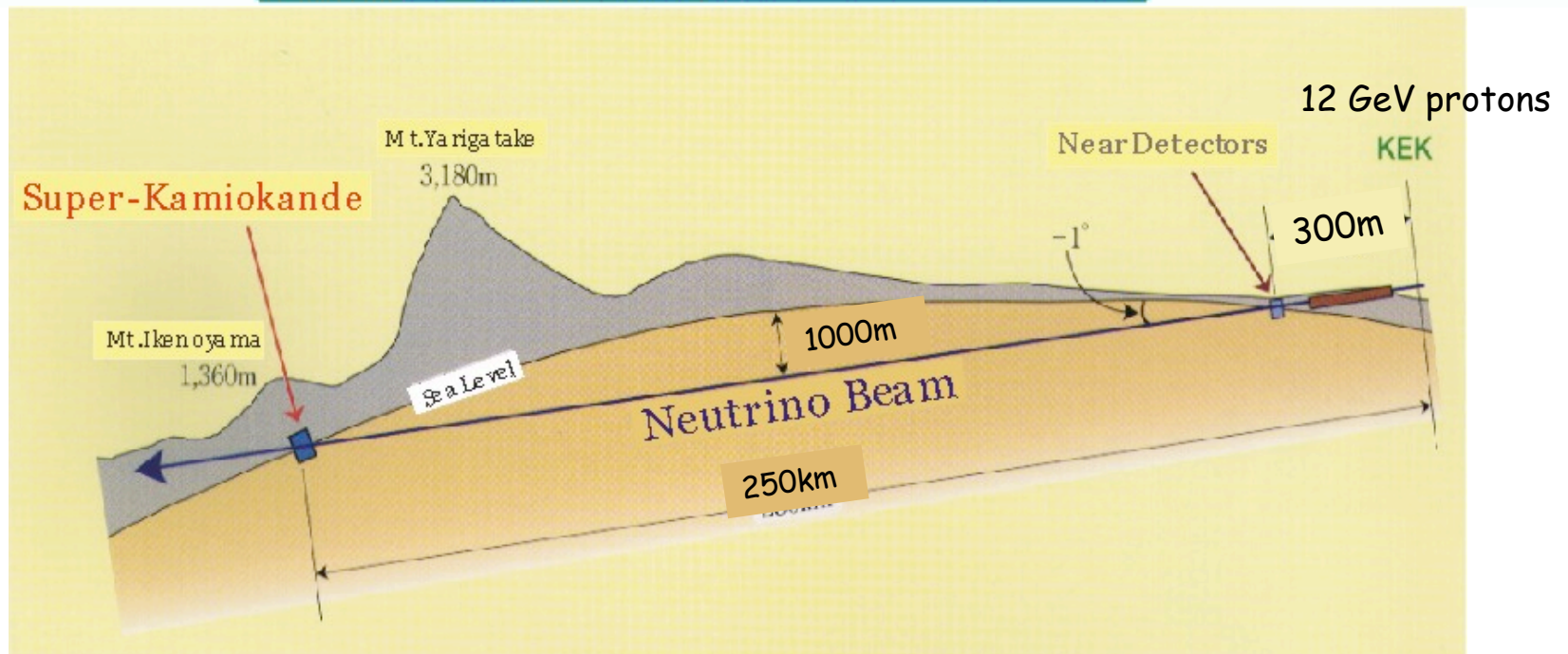
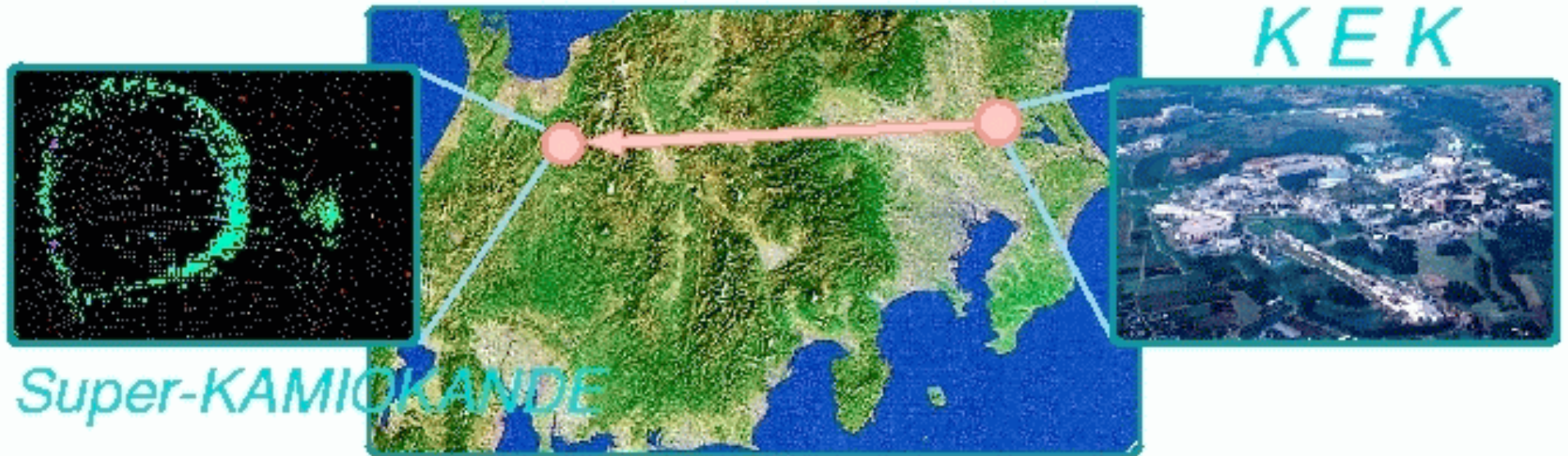


$L_\nu / E_\nu$  distribution



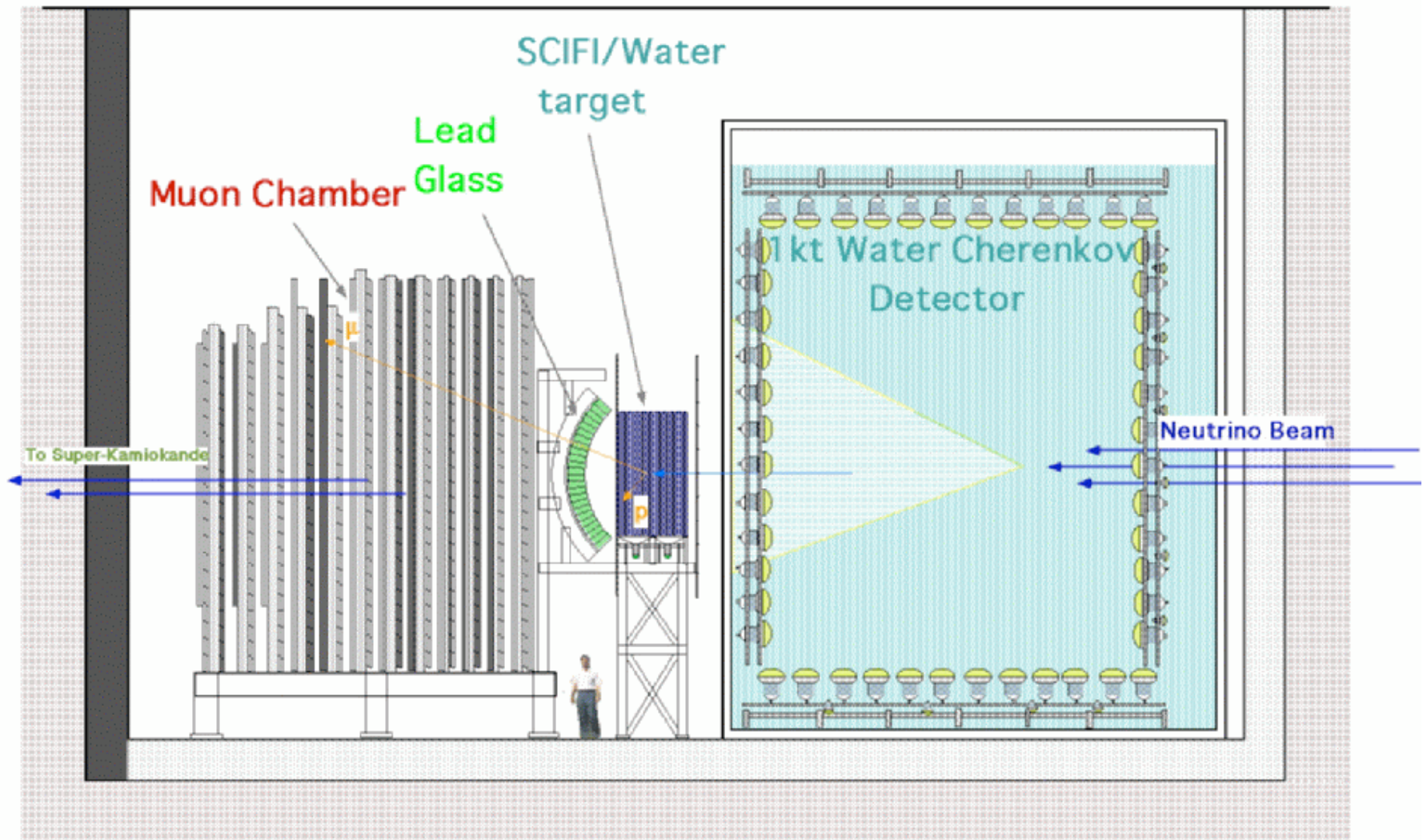
# Verifying The Atmospheric Neutrino Results With Artificial Muon-neutrino Beams

# K2K: Long-baseline Neutrino Oscillation Experiment

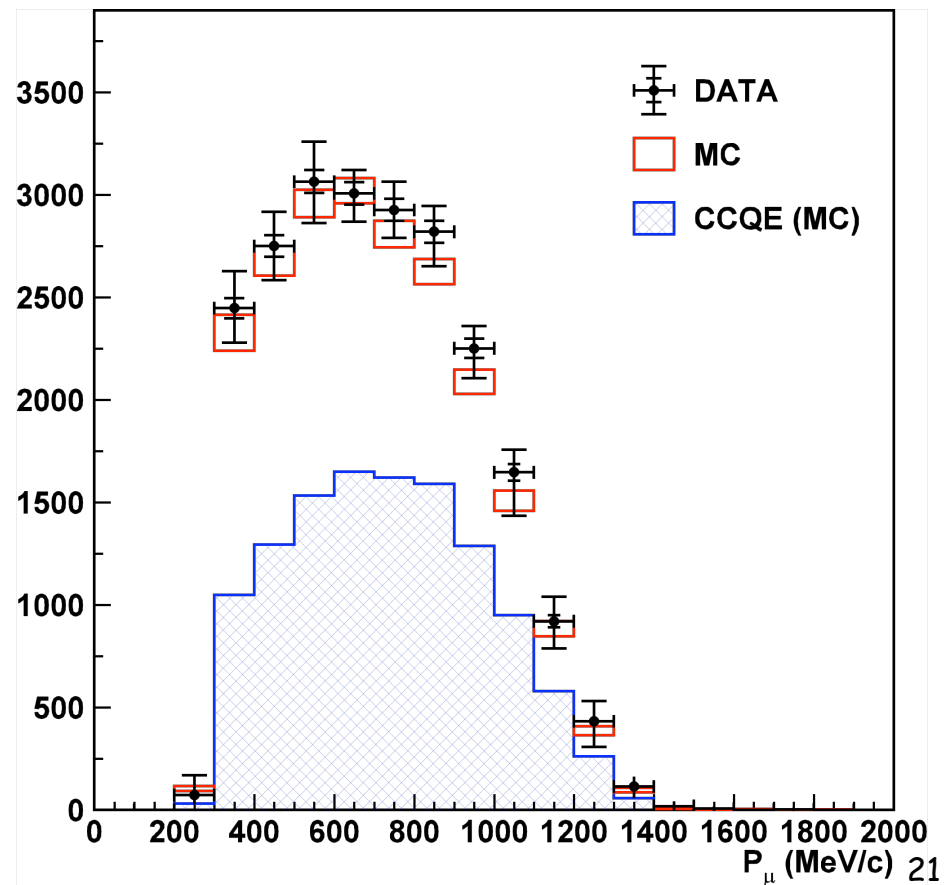
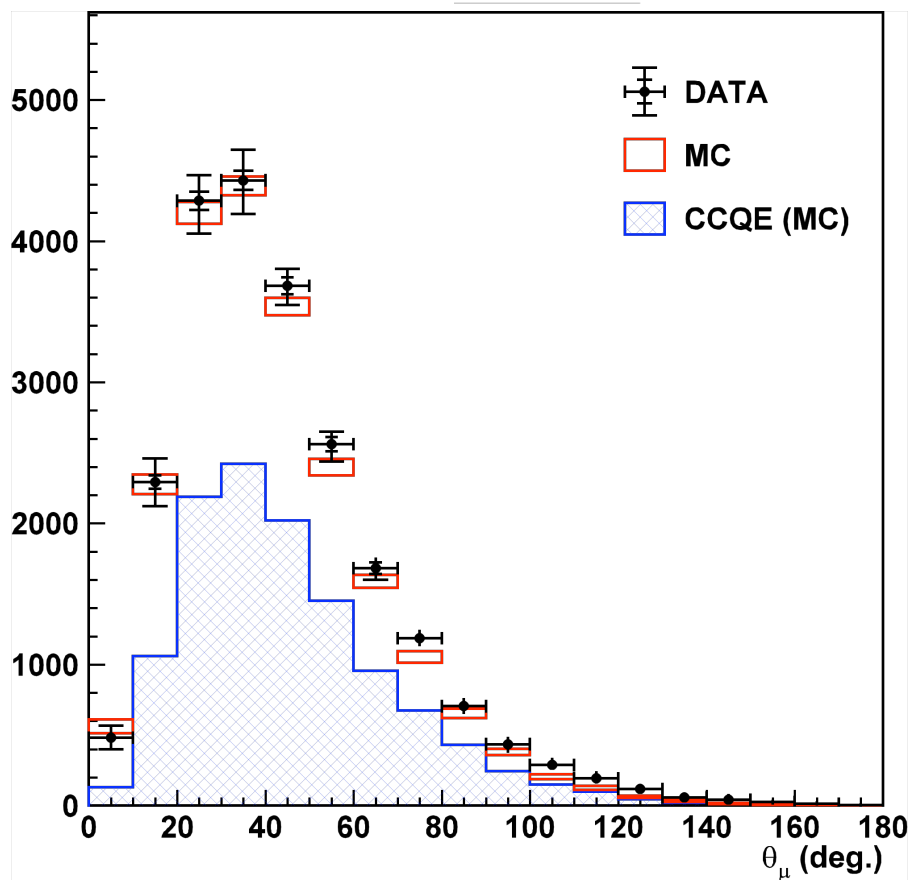
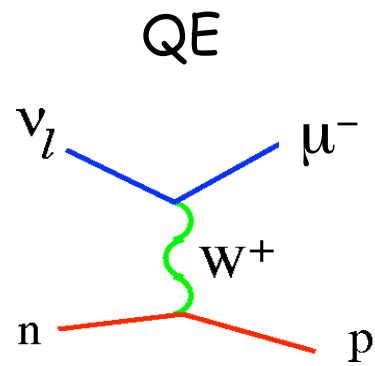
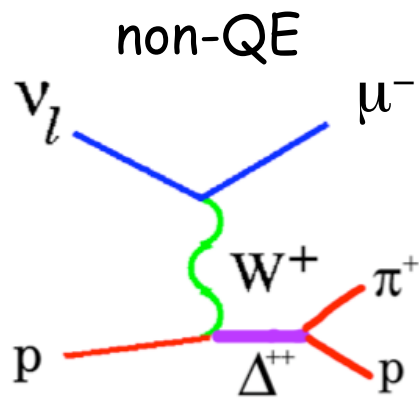


# Near Detector of K2K

- To determine the neutrino flux and its energy distribution
- Study low-energy charged-current neutrino reactions

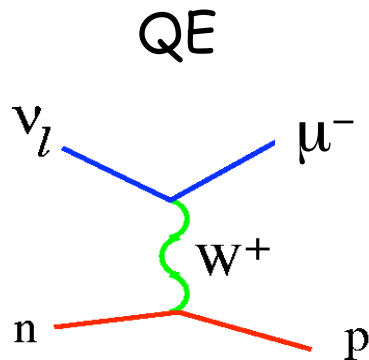


# Study Charged-current With Near Detector



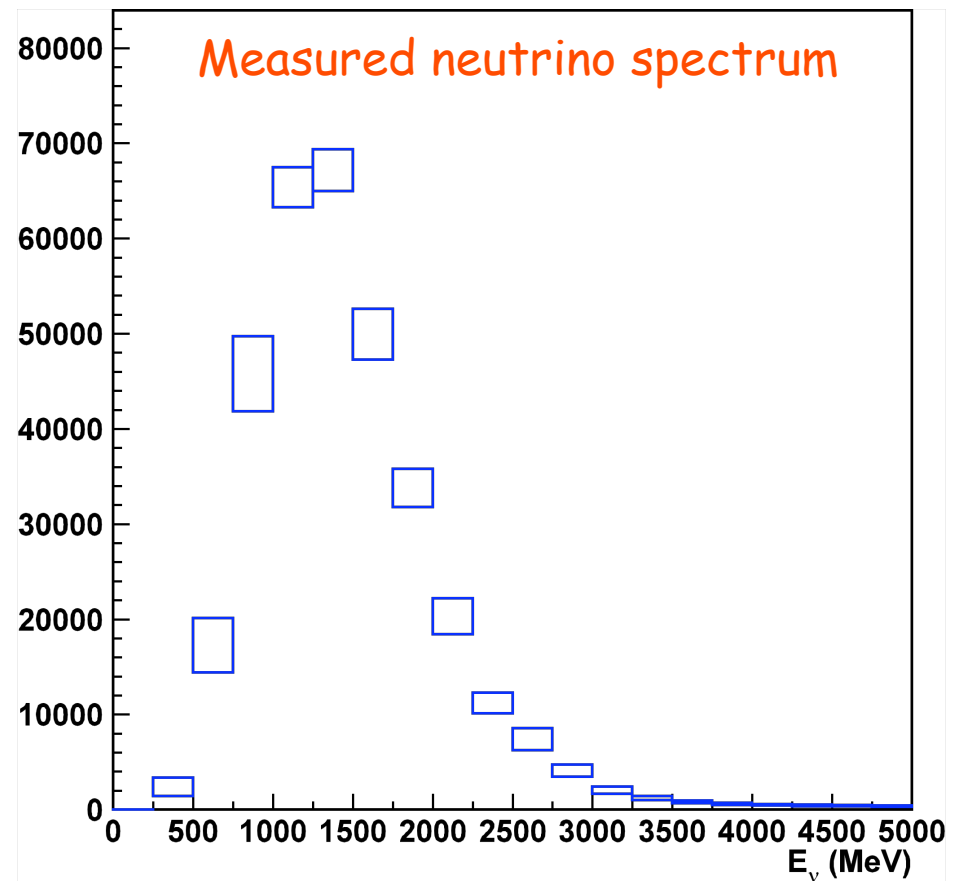
# Determining The Neutrino Spectrum

The energy of  $\nu_\mu$  is related to the muon momentum in quasi-elastic scattering:

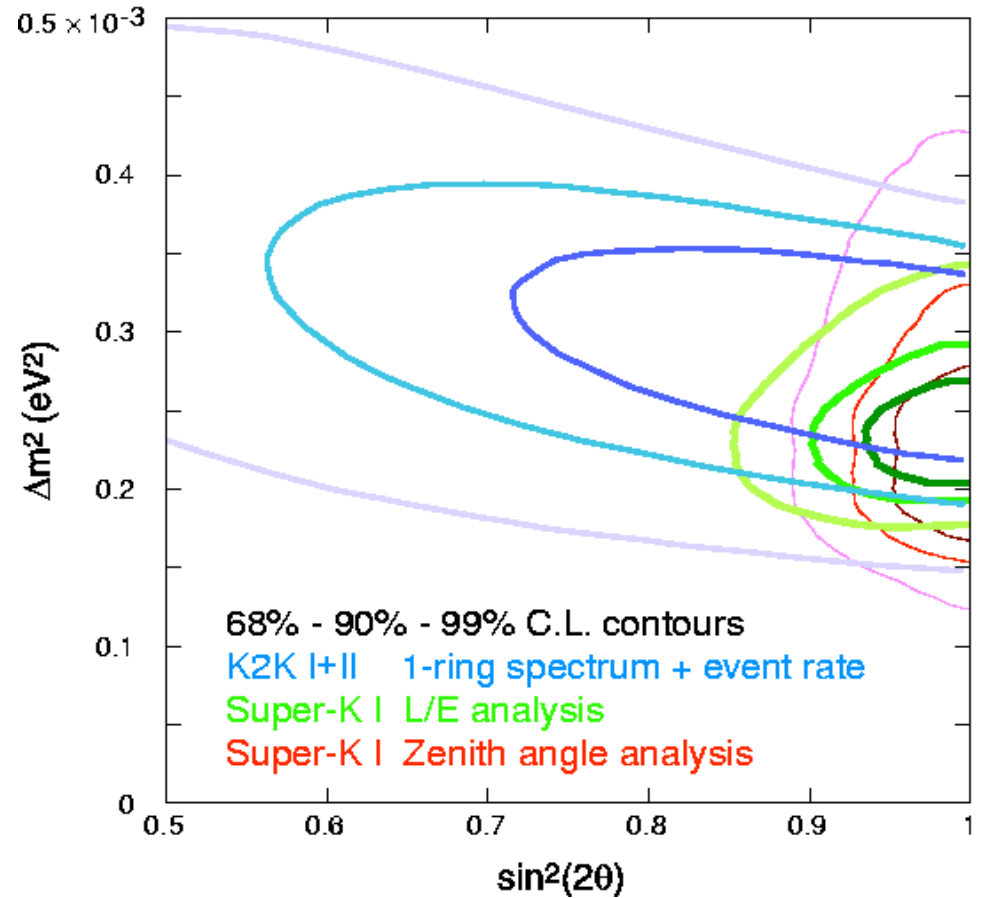
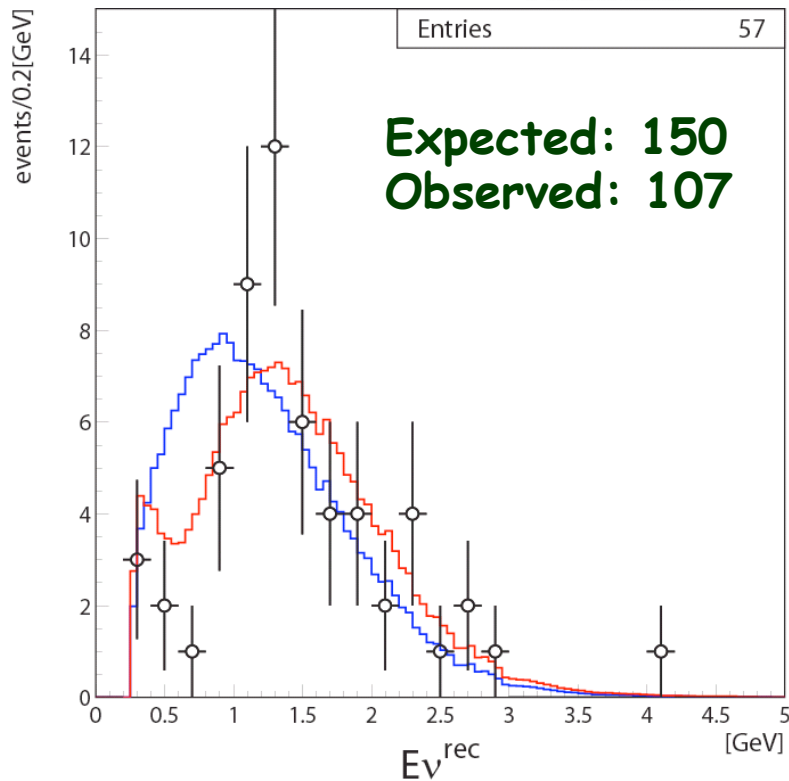
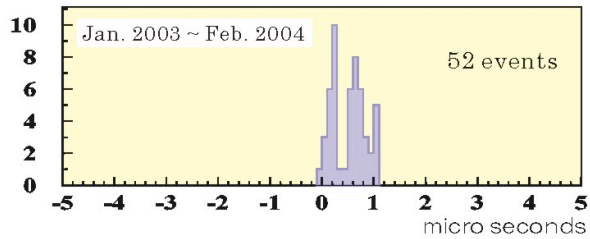
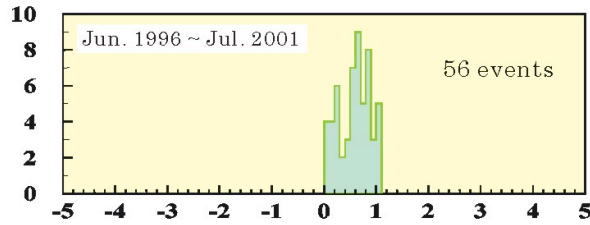


From kinematics:

$$E_\nu = \frac{m_N E_\mu - \frac{1}{2} m_\mu^2}{m_N - E_\mu + p_\mu \cos \theta_\mu}$$



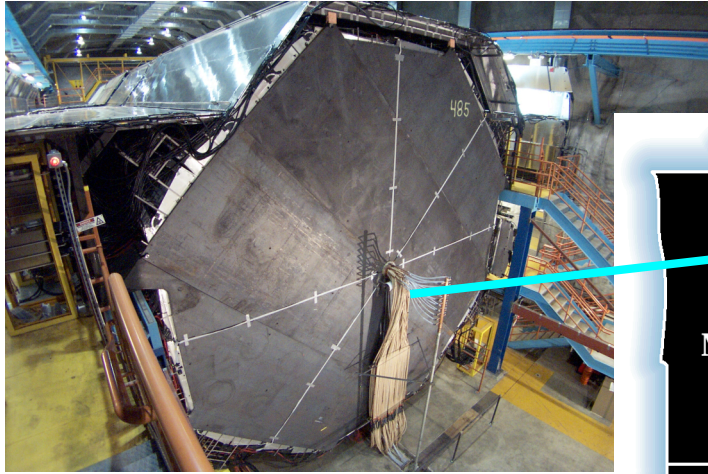
# Indeed $\nu_\mu$ Oscillates !!



Best fit result:

$$\Delta m^2 = 2.8 \times 10^{-3} \text{ eV}^2$$

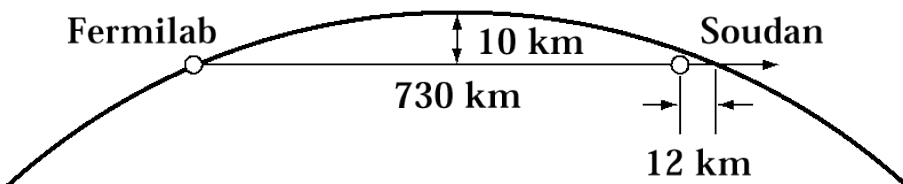
$$\sin^2 2\theta = 1.0$$



# MINOS

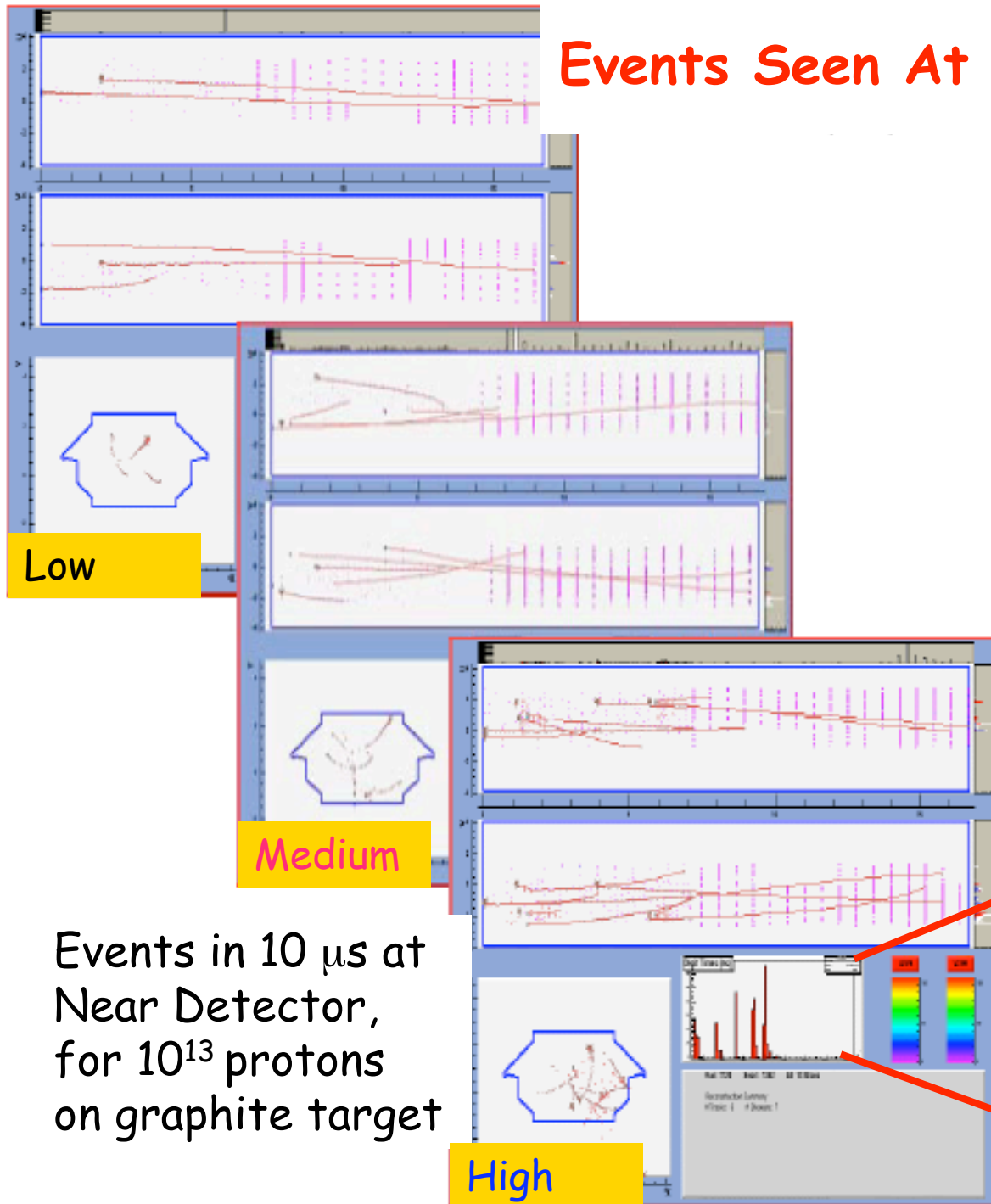


5.4 kt magnetized iron-scintillator used as a tracker and calorimeter

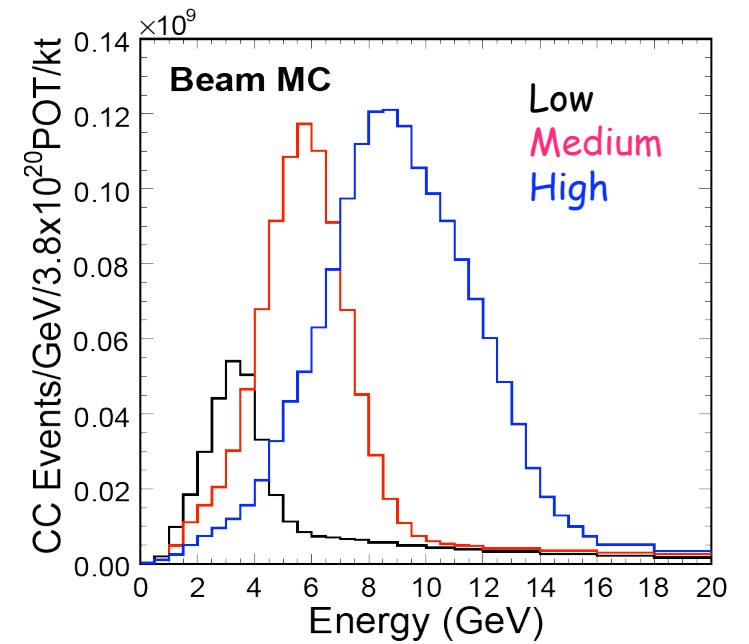




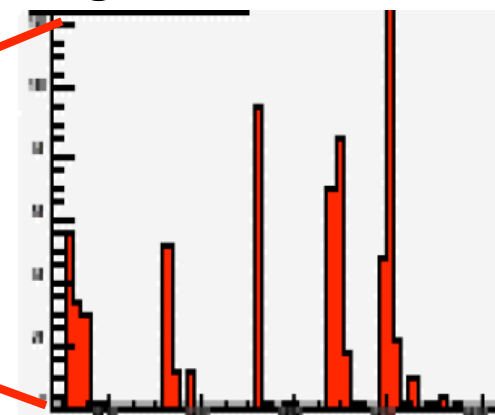
# Events Seen At MINOS Near Detector



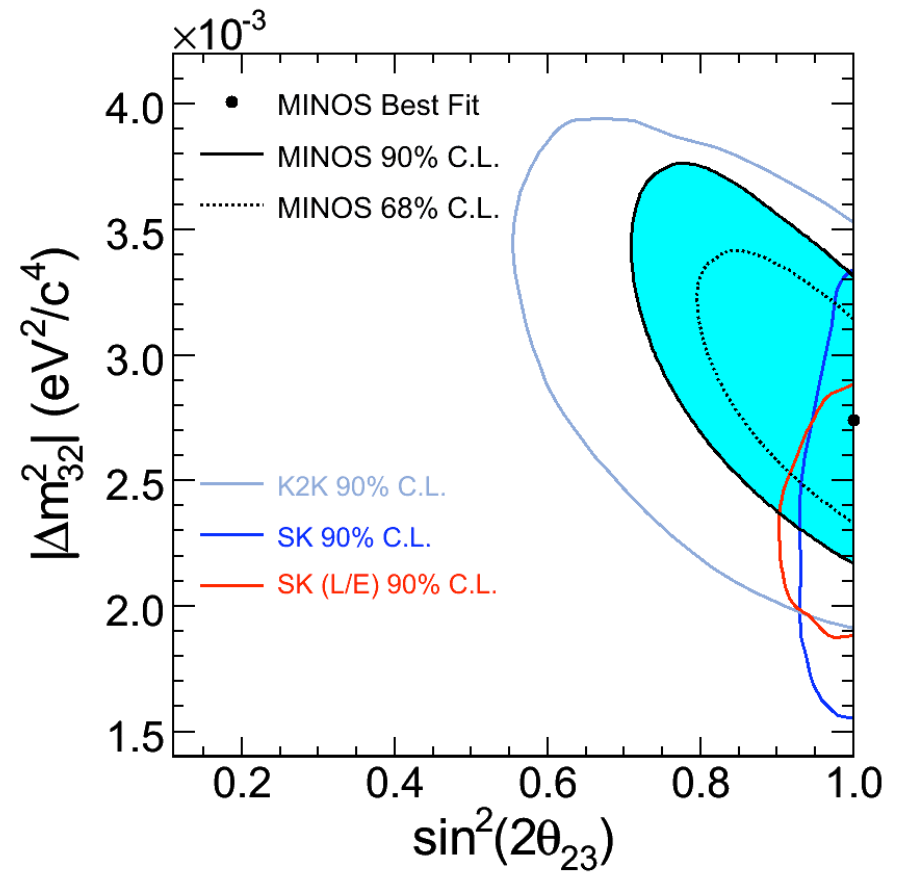
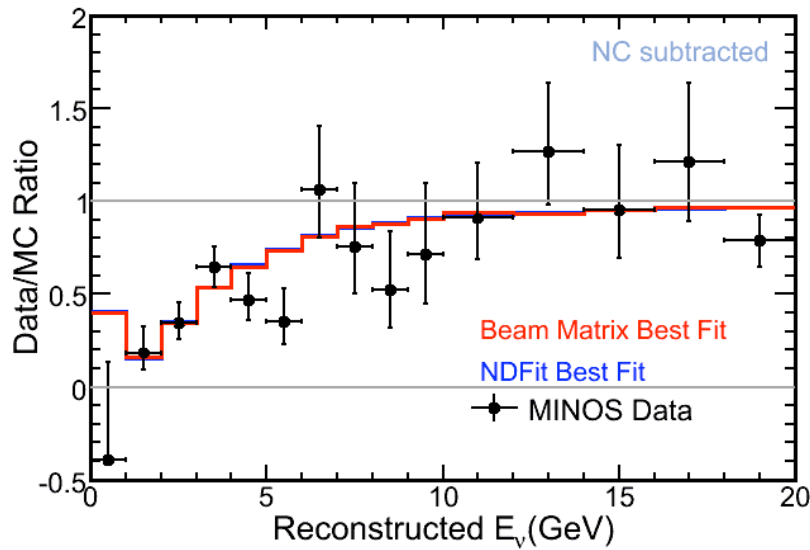
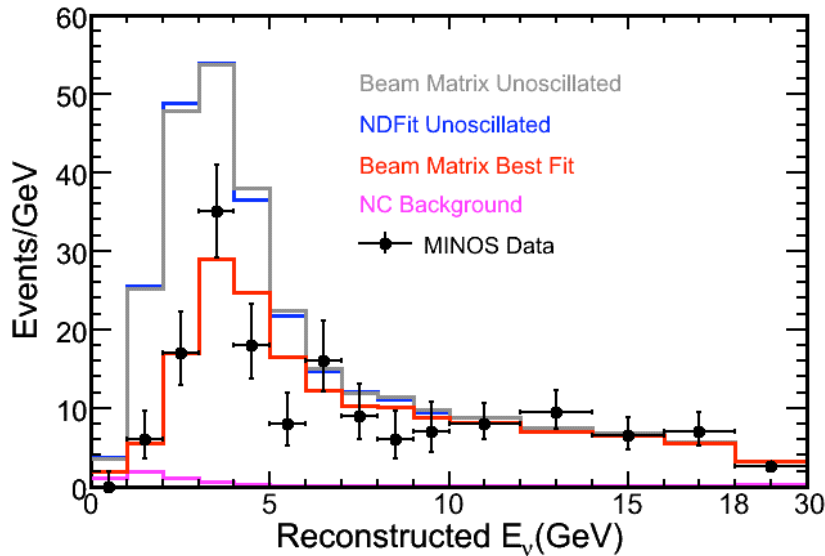
Events in  $10 \mu\text{s}$  at Near Detector, for  $10^{13}$  protons on graphite target



Signal Time (nsec)



# First Disappearance Result From MINOS



$$|\Delta m_{32}^2| = 2.74^{+0.44}_{-0.26} \text{ (stat + syst)} \times 10^{-3} \text{ eV}^2$$

$$\sin^2 2\theta_{23} = 1.00_{-0.13} \text{ (stat + syst)}$$

## Summary

- Study of atmospheric neutrinos led to the discovery of significant disappearance of  $\nu_{\mu}$  but not  $\nu_e$ , providing strong evidence for another type of neutrino oscillation.
- This observation has been confirmed by other underground, and accelerator-based experiments.
- Again, the new set of mixing parameters  $\theta$  and  $\Delta m^2$  using the 2-flavour formalism is measured.