

Lepton Flavour Violation in SUSY GUTs

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November 7, 2006

*'TeV scale Physics at the LHC', CCAST Beijing,
Sept 15th 2006*

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U_{e3} and $\text{BR}(\mu \rightarrow e\gamma)$ in SUSY SO(10)

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Outline

- Introduction to $SO(10)$
- Two Simple examples
- Invariance of the Casas-Ibara R matrix
- U_{e3} versus $\mu \rightarrow e\gamma$
- Summary

General SO(10)

Each SM family, plus a right-handed neutrino in a single 16-dim rep.

$$\mathbf{W}_{\text{SO}(10)}^{\text{Y}} = \mathbf{h}_{ij}^{10} \mathbf{16}_i \mathbf{16}_j \mathbf{10}_H + \mathbf{h}_{ij}^{126} \mathbf{16}_i \mathbf{16}_j \mathbf{126}_H + \mathbf{h}_{ij}^{120} \mathbf{16}_i \mathbf{16}_j \mathbf{120}_H \quad (1)$$

Here the matrices $h^{10,126}$ are symmetric and h^{120} is anti-symmetric.

$$\begin{aligned} 16 \ 16 \ 10 &\supset 5(uu^c + \nu\nu^c) + \bar{5}(dd^c + ee^c) \\ 16 \ 16 \ 126 &\supset 1\nu^c\nu^c + 15\nu\nu + 5(uu^c - 3\nu\nu^c) + 4\bar{5}(dd^c - 3ee^c) \\ 16 \ 16 \ 120 &\supset 5\nu\nu^c + 45uu^c + \bar{5}(dd^c + ee^c) + 4\bar{5}(dd^c - 3ee^c) \end{aligned} \quad (2)$$

The resulting tree level mass matrices are as follows,

$$M^u = M_{10}^5 + M_{126}^5 + M_{120}^{45} \quad (3)$$

$$M^\nu = M_{10}^5 - 3M_{126}^5 + M_{120}^5 \quad (4)$$

$$M^d = M_{10}^{\bar{5}} + M_{126}^{\bar{45}} + M_{120}^{\bar{5}} + M_{120}^{\bar{45}} \quad (5)$$

$$M^e = M_{10}^{\bar{5}} - 3M_{126}^{\bar{45}} + M_{120}^{\bar{5}} - 3M_{120}^{\bar{45}} \quad (6)$$

$$M_{LL}^\nu = M_{126}^{15} \quad (7)$$

$$M_R^\nu = M_{126}^1 \quad (8)$$

General SO(10)

Each SM family, plus a right-handed neutrino in a single 16-dim rep.

$$\mathbf{W}_{\text{SO}(10)}^Y = \mathbf{h}_{ij}^{10} \mathbf{16}_i \mathbf{16}_j \mathbf{10}_H + \mathbf{h}_{ij}^{126} \mathbf{16}_i \mathbf{16}_j \mathbf{126}_H + \mathbf{h}_{ij}^{120} \mathbf{16}_i \mathbf{16}_j \mathbf{120}_H \quad (9)$$

Here the matrices $h^{10,126}$ are symmetric and h^{120} is anti-symmetric.

$$\begin{aligned} \mathbf{16} \mathbf{16} \mathbf{10} &\supset \mathbf{5}(\mathbf{uu}^c + \nu\nu^c) + \bar{\mathbf{5}}(\mathbf{dd}^c + \mathbf{ee}^c) \\ \mathbf{16} \mathbf{16} \mathbf{126} &\supset \mathbf{1}\nu^c\nu^c + \mathbf{15}\nu\nu + \mathbf{5}(\mathbf{uu}^c - \mathbf{3}\nu\nu^c) + \mathbf{4}\bar{\mathbf{5}}(\mathbf{dd}^c - \mathbf{3ee}^c) \\ \mathbf{16} \mathbf{16} \mathbf{120} &\supset \mathbf{5}\nu\nu^c + \mathbf{45}\mathbf{uu}^c + \bar{\mathbf{5}}(\mathbf{dd}^c + \mathbf{ee}^c) + \mathbf{4}\bar{\mathbf{5}}(\mathbf{dd}^c - \mathbf{3ee}^c) \end{aligned} \quad (10)$$

The resulting tree level mass matrices are as follows,

$$M^u = M_{10}^5 + M_{126}^5 + M_{120}^{45} \quad (11)$$

$$M^\nu = M_{10}^5 - 3M_{126}^5 + M_{120}^5 \quad (12)$$

$$M^d = M_{10}^{\bar{5}} + M_{126}^{\bar{45}} + M_{120}^{\bar{5}} + M_{120}^{\bar{45}} \quad (13)$$

$$M^e = M_{10}^{\bar{5}} - 3M_{126}^{\bar{45}} + M_{120}^{\bar{5}} - 3M_{120}^{\bar{45}} \quad (14)$$

$$M_{LL}^\nu = M_{126}^{15} \quad (15)$$

$$M_R^\nu = M_{126}^1 \quad (16)$$

General SO(10)

Each SM family, plus a right-handed neutrino in a single 16-dim rep.

$$\mathbf{W}_{\text{SO}(10)}^Y = \mathbf{h}_{ij}^{10} \mathbf{16}_i \mathbf{16}_j \mathbf{10}_H + \mathbf{h}_{ij}^{126} \mathbf{16}_i \mathbf{16}_j \mathbf{126}_H + \mathbf{h}_{ij}^{120} \mathbf{16}_i \mathbf{16}_j \mathbf{120}_H \quad (17)$$

Here the matrices $h^{10,126}$ are symmetric and h^{120} is anti-symmetric.

$$\begin{aligned} \mathbf{16} \mathbf{16} \mathbf{10} &\supset \mathbf{5}(\mathbf{uu}^c + \nu\nu^c) + \bar{\mathbf{5}}(\mathbf{dd}^c + \mathbf{ee}^c) \\ \mathbf{16} \mathbf{16} \mathbf{126} &\supset \mathbf{1}\nu^c\nu^c + \mathbf{15}\nu\nu + \mathbf{5}(\mathbf{uu}^c - \mathbf{3}\nu\nu^c) + \mathbf{4}\bar{\mathbf{5}}(\mathbf{dd}^c - \mathbf{3ee}^c) \\ \mathbf{16} \mathbf{16} \mathbf{120} &\supset \mathbf{5}\nu\nu^c + \mathbf{45}\mathbf{uu}^c + \bar{\mathbf{5}}(\mathbf{dd}^c + \mathbf{ee}^c) + \mathbf{4}\bar{\mathbf{5}}(\mathbf{dd}^c - \mathbf{3ee}^c) \end{aligned} \quad (18)$$

The resulting tree level mass matrices are as follows,

$$M^u = M_{10}^5 + M_{126}^5 + M_{120}^{45} \quad (19)$$

$$M^\nu = M_{10}^5 - 3M_{126}^5 + M_{120}^5 \quad (20)$$

$$M^d = M_{10}^{\bar{5}} + M_{126}^{\bar{45}} + M_{120}^{\bar{5}} + M_{120}^{\bar{45}} \quad (21)$$

$$M^e = M_{10}^{\bar{5}} - 3M_{126}^{\bar{45}} + M_{120}^{\bar{5}} - 3M_{120}^{\bar{45}} \quad (22)$$

$$M_{LL}^\nu = M_{126}^{15} \quad (23)$$

$$M_R^\nu = M_{126}^1 \quad (24)$$

General SO(10)

Each SM family, plus a right-handed neutrino in a single 16-dim rep.

$$\mathbf{W}_{\text{SO}(10)}^Y = \mathbf{h}_{ij}^{10} \mathbf{16}_i \mathbf{16}_j \mathbf{10}_H + \mathbf{h}_{ij}^{126} \mathbf{16}_i \mathbf{16}_j \mathbf{126}_H + \mathbf{h}_{ij}^{120} \mathbf{16}_i \mathbf{16}_j \mathbf{120}_H \quad (25)$$

Here the matrices $h^{10,126}$ are symmetric and h^{120} is anti-symmetric.

$$\begin{aligned} \mathbf{16} \mathbf{16} \mathbf{10} &\supset \mathbf{5}(\mathbf{uu}^c + \nu\nu^c) + \bar{\mathbf{5}}(\mathbf{dd}^c + \mathbf{ee}^c) \\ \mathbf{16} \mathbf{16} \mathbf{126} &\supset \mathbf{1}\nu^c\nu^c + \mathbf{15}\nu\nu + \mathbf{5}(\mathbf{uu}^c - \mathbf{3}\nu\nu^c) + \mathbf{4}\bar{\mathbf{5}}(\mathbf{dd}^c - \mathbf{3ee}^c) \\ \mathbf{16} \mathbf{16} \mathbf{120} &\supset \mathbf{5}\nu\nu^c + \mathbf{45}\mathbf{uu}^c + \bar{\mathbf{5}}(\mathbf{dd}^c + \mathbf{ee}^c) + \mathbf{4}\bar{\mathbf{5}}(\mathbf{dd}^c - \mathbf{3ee}^c) \end{aligned} \quad (26)$$

The resulting tree level mass matrices are as follows,

$$M^u = M_{10}^5 + M_{126}^5 + M_{120}^{45} \quad (27)$$

$$M^\nu = M_{10}^5 - 3M_{126}^5 + M_{120}^5 \quad (28)$$

$$M^d = M_{10}^{\bar{5}} + M_{126}^{\bar{45}} + M_{120}^{\bar{5}} + M_{120}^{\bar{45}} \quad (29)$$

$$M^e = M_{10}^{\bar{5}} - 3M_{126}^{\bar{45}} + M_{120}^{\bar{5}} - 3M_{120}^{\bar{45}} \quad (30)$$

$$M_{LL}^\nu = M_{126}^{15} \quad (31)$$

$$M_R^\nu = M_{126}^1 \quad (32)$$

Two simple models of Y_ν mixing

The size of mixing in Y_ν determines LFV rates

Masiero et.al. hep-ph/0209303

- **Case I:** CKM type mixing in $Y_\nu \implies$ small LFV rates
 - $\tan\beta = 5$ and $\tan\beta = 50$
- **Case II:** MNS type mixing in $Y_\nu \implies$ large LFV rates
 - $\tan\beta = 5$ and $\tan\beta = 50$

I: CKM type mixings in Y_ν

Introduce separate Higgs couplings for the up and down type fermions.

$$\mathbf{W}_{\text{SO}(10)} = \frac{1}{2} \mathbf{h}_{ij}^{u,\nu} \mathbf{16}_i \mathbf{16}_j \mathbf{10}_u + \frac{1}{2} \mathbf{h}_{ij}^{d,e} \mathbf{16}_i \mathbf{16}_j \mathbf{10}_d + \frac{1}{2} \mathbf{h}_{ij}^R \mathbf{16}_i \mathbf{16}_j \mathbf{126} \quad (33)$$

Leading to the $MSSM + \nu_R$ superpotential,

$$\mathbf{W} = \mathbf{h}_{ij}^u \mathbf{Q}_i \mathbf{u}_j \mathbf{H}_u + \mathbf{h}_{ij}^\nu \mathbf{L}_i \nu_j \mathbf{H}_u + \mathbf{h}_{ij}^d \mathbf{Q}_i \mathbf{d}_j \mathbf{H}_d + \mathbf{h}_{ij}^e \mathbf{L}_i \mathbf{e}_j \mathbf{H}_d + \frac{1}{2} \mathbf{M}_{ij}^R \nu_i \nu_j \quad (34)$$

Here, $h^u = h^\nu$ and $h^d = h^e$. Here the symmetric h^u is diagonalised by,

$$\mathbf{h}_{\text{diag}}^u = \mathbf{V}_{\text{CKM}} \mathbf{h}^u \mathbf{V}_{\text{CKM}}^T \quad (35)$$

so that,

$$\mathbf{h}^\nu = \mathbf{V}_{\text{CKM}}^T \mathbf{h}_{\text{diag}}^u \mathbf{V}_{\text{CKM}} \quad (36)$$

The type I seesaw mechanism then gives,

$$\mathbf{m}_\nu = \mathbf{h}^\nu \mathbf{M}_R^{-1} \mathbf{h}^{\nu T} \mathbf{v}_u^2 \quad (37)$$

So that the Majorana Mass matrix is of the form,

$$\mathbf{M}_R = \mathbf{v}_u^2 \mathbf{V}_{\text{CKM}}^T \mathbf{h}_{\text{diag}}^u \mathbf{V}_{\text{CKM}} \mathbf{U}_{\text{MNS}} (\mathbf{m}_\nu^{\text{diag}})^{-1} \mathbf{U}_{\text{MNS}}^T \mathbf{V}_{\text{CKM}}^T \mathbf{h}_{\text{diag}}^u \mathbf{V}_{\text{CKM}} \quad (38)$$

II: MNS type mixings in Y_ν

Introduce separate Higgs couplings for the up and down type fermions. To get MNS like mixing we must introduce an asymmetrical Higgs coupling, Φ .

$$\mathbf{W}_{\text{SO}(10)} = \frac{1}{2} \mathbf{h}_{ii}^{u,\nu} 16_i 16_i 10_u + \frac{1}{2} \mathbf{h}_{ij}^{d,e} 16_i 16_j \Phi + \frac{1}{2} \mathbf{h}_{ij}^R 16_i 16_j 126 \quad (39)$$

If we write it in $\text{SU}(5)$ language then we have,

$$\mathbf{W}_{\text{SU}(5)} = \frac{1}{2} \mathbf{h}_{ii}^u 10_i 10_i 5_u + \mathbf{h}_{ii}^\nu \bar{5}_i 1_i 5_u + \mathbf{h}_{ij}^d 10_i \bar{5}_j \bar{5}_d + \frac{1}{2} \mathbf{M}_{ij}^R 1_i 1_j \quad (40)$$

Here, $h_{ii}^u = h_{ii}^\nu$. Rotating to the diagonal down quark basis,

$$\mathbf{h}_{\text{diag}}^d = \mathbf{V}_{\text{CKM}}^T \mathbf{h}^d \mathbf{U}_{\text{MNS}}^T \quad (41)$$

And,

$$h^u = V_{\text{CKM}}^T h_{\text{diag}}^u V_{\text{CKM}}^T \quad (42)$$

$$\mathbf{h}^\nu = \mathbf{U}_{\text{MNS}} \mathbf{h}_{\text{diag}}^u \quad (43)$$

The type I seesaw mechanism then gives,

$$\mathbf{m}_\nu = \mathbf{h}^\nu \mathbf{M}_R^{-1} \mathbf{h}^{\nu T} \mathbf{v}_u^2 \quad (44)$$

Inverting this relation we have the right-handed neutrino mass,

$$\mathbf{M}_R = \mathbf{h}^{\nu T} \mathbf{m}_\nu^{-1} \mathbf{h}^\nu \mathbf{v}_u^2 = \mathbf{h}_{\text{diag}}^u (\mathbf{m}_\nu^{\text{diag}})^{-1} \mathbf{h}_{\text{diag}}^u \mathbf{v}_u^2 = \text{diag} \left\{ \frac{\mathbf{m}_u^2}{\mathbf{m}_{\nu_1}}, \frac{\mathbf{m}_c^2}{\mathbf{m}_{\nu_2}}, \frac{\mathbf{m}_t^2}{\mathbf{m}_{\nu_3}} \right\} \quad (45)$$

The Casas-Ibara R matrix

Usually defined in the ‘Flavour Basis’; **diagonal** M_{RR}, Y_{LR}^E

$$\mathcal{L} = -Y_{LR}^{E \text{ diag}} H_d \bar{L} E_R - Y_{LR}^\nu H_u \bar{L} \nu_R + \frac{1}{2} \nu_R^T M_{RR}^{\text{diag}} \nu_R \quad (46)$$

The seesaw mass is,

$$m_\nu = v_u^2 Y_{LR}^\nu M_{RR}^{\text{diag}^{-1}} Y_{LR}^{\nu T} \quad (47)$$

Which is diagonalised by the MNS matrix,

$$m_\nu^{\text{diag}} = U_{\text{MNS}}^\dagger m_\nu U_{\text{MNS}}^* \quad (48)$$

Inverting this equation gives you,

$$\mathbf{1} = v_u^2 (m_\nu^{\text{diag}})^{-1/2} U_{\text{MNS}}^\dagger Y_{LR}^\nu M_{RR}^{\text{diag}^{-1/2}} M_{RR}^{\text{diag}^{-1/2}} Y_{LR}^{\nu T} U_{\text{MNS}}^* (m_\nu^{\text{diag}})^{-1/2} \quad (49)$$

$$= \mathbf{R} \mathbf{R}^T \quad (50)$$

So that,

$$\mathbf{R} = v_u (m_\nu^{\text{diag}})^{-1/2} U_{\text{MNS}}^\dagger Y_{LR}^\nu M_{RR}^{\text{diag}^{-1/2}} \quad (51)$$

Is the R matrix of Casas-Ibara, hep-ph/0103065, defined in the flavour basis.

$$Y_{LR}^\nu = U_{\text{MNS}} (m_\nu^{\text{diag}})^{1/2} \mathbf{R} M_{RR}^{\text{diag}^{1/2}} \mathbf{1}/v_u \quad (52)$$

Invariance of the Casas-Ibara R matrix

Recently S.F.King hep-ph/0610239, showed that R is basis invariant.

All models with the same, R matrix, are related by a simple basis rotation.

i.e. Y_{LR}^E , Y_{LR}^ν , M_{RR} and R defines a ‘class’ of models:

In a general basis we have the seesaw mass,

$$\mathbf{m}_\nu = \mathbf{v}_u^2 \mathbf{Y}_{LR}^\nu \mathbf{M}_{RR}^{\text{diag}^{-1}} \mathbf{Y}_{LR}^{\nu T} \quad (53)$$

Inverting again gives,

$$\mathbf{1} = \mathbf{v}_u^2 \mathbf{m}_{\sqrt{\nu}}^{-1} \mathbf{Y}_{LR}^\nu \sqrt{\mathbf{M}_{RR}^{-1}} \sqrt{\mathbf{M}_{RR}^{-1}}^T \mathbf{Y}_{LR}^{\nu T} (\mathbf{m}_{\sqrt{\nu}}^{-1})^T = \mathbf{R} \mathbf{R}^T \quad (54)$$

Where, $\mathbf{M}_{RR}^{-1} = \sqrt{\mathbf{M}_{RR}^{-1}} \sqrt{\mathbf{M}_{RR}^{-1}}^T$ and $\mathbf{m}_\nu = \mathbf{m}_{\sqrt{\nu}} \mathbf{m}_{\sqrt{\nu}}^T$

So that the R matrix in a general basis is,

$$\mathbf{R} = \mathbf{v}_u^2 \mathbf{m}_{\sqrt{\nu}}^{-1} \mathbf{Y}_{LR}^\nu \sqrt{\mathbf{M}_{RR}^{-1}} \quad (55)$$

And R is clearly basis invariant under a basis rotation,

$$\mathbf{m}_{\sqrt{\nu}}^{-1} \rightarrow \mathbf{m}_{\sqrt{\nu}}^{-1} \mathbf{V}_{E_L}^\dagger \quad (56)$$

$$\sqrt{\mathbf{M}_{RR}^{-1}} \rightarrow \mathbf{V}_{\nu_R} \sqrt{\mathbf{M}_{RR}^{-1}} \quad (57)$$

$$\mathbf{Y}_{LR}^\nu \rightarrow \mathbf{V}_{E_L} \mathbf{Y}_{LR}^\nu \mathbf{V}_{\nu_R}^\dagger \quad (58)$$

R for our two simple models

- CKM mixing Case:

$$\mathbf{R} = \mathbf{1}$$

- MNS mixing Case:

$$\mathbf{R} = \mathbf{1}$$

The two models are related by a non-unitary transformation of the right handed neutrino field,

$$\mathbf{S}^{-1} = \mathbf{V}_{\text{CKM}}^{\text{T}} \mathbf{h}_{\text{diag}}^{\text{u}-1} \mathbf{V}_{\text{CKM}} \mathbf{U}_{\text{MNS}} \mathbf{h}_{\text{diag}}^{\text{u}} \quad (59)$$

With R also invariant under,

$$\mathbf{Y}_{\text{LR}}^{\nu} \rightarrow \mathbf{Y}_{\text{LR}}^{\nu} \mathbf{S}^{-1} \quad (60)$$

$$\mathbf{M}_{\text{R}} \rightarrow \mathbf{S}^{\text{T}-1} \mathbf{M}_{\text{R}} \mathbf{S}^{-1} \quad (61)$$

$\mu \rightarrow e\gamma$ in model II

Approximate off-diagonal slepton mass terms induced by RG running as,

$$[\mathbf{m}_{\tilde{l}}^2]_{21} \approx -\frac{1}{8\pi^2}(3\mathbf{m}_0^2 + \mathbf{A}_0^2) [\mathbf{h}^\nu \mathbf{h}^{\nu T}]_{21} \ln \left(\frac{\mathbf{M}_{\text{GUT}}}{\mathbf{M}_{\text{R}}} \right) \quad (62)$$

Then the branching ratio for $l_i \rightarrow l_j \gamma$ is approximately,

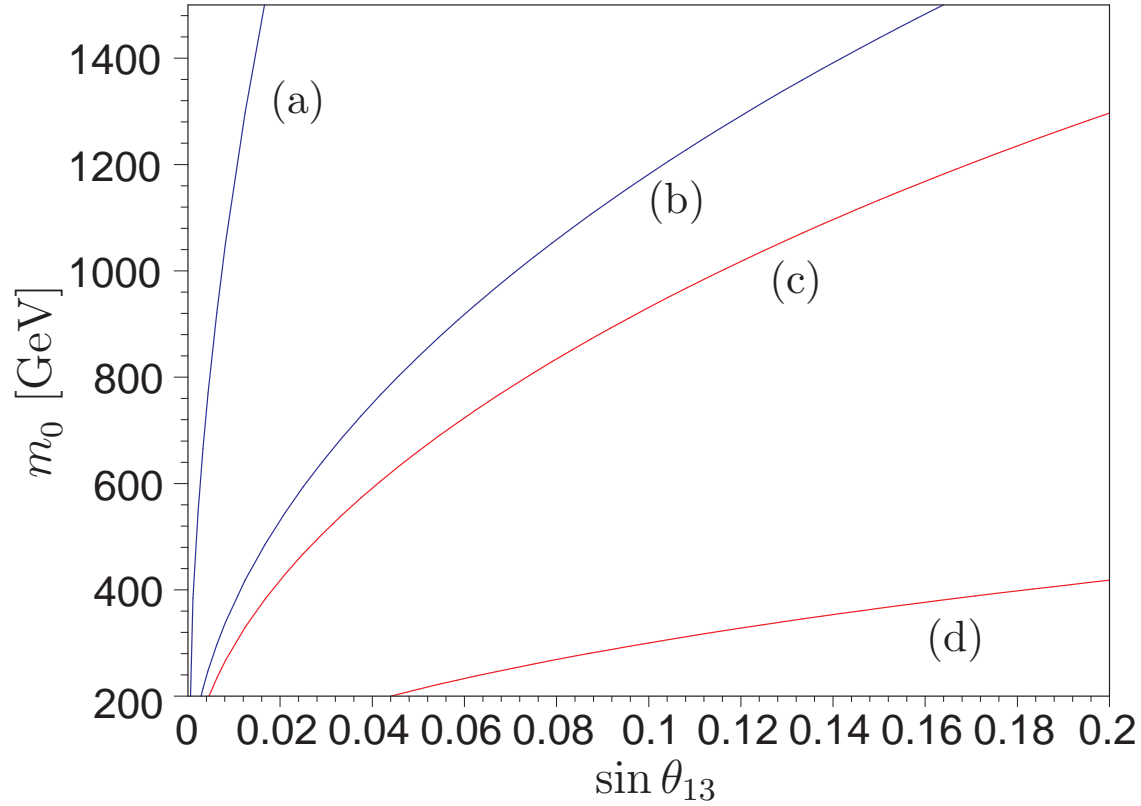
$$\text{Br}(l_i \rightarrow l_j \gamma) \approx \frac{\alpha^2}{\mathbf{G}_{\text{F}}^2} \frac{([\mathbf{m}_{\tilde{l}}^2]_{21})^2}{\mathbf{m}_{\text{SUSY}}^8} \tan^2 \beta \leq \mathbf{B} \quad (63)$$

Here B is the experimental bound set on this LFV decay. The present value is $B = 1.2 \times 10^{-11}$ at 90% C.L., but in the near future it could become $B = 5 \times 10^{-14}$ at 90% C.L.

For the MNS mixing model we have,

$$[\mathbf{h}^\nu \mathbf{h}^{\nu T}]_{21} = \mathbf{y}_{\text{u}}^2 \mathbf{U}_{\mu 1} \mathbf{U}_{e 1} + \mathbf{y}_{\text{c}}^2 \mathbf{U}_{\mu 2} \mathbf{U}_{e 2} + \mathbf{y}_{\text{t}}^2 \mathbf{U}_{\mu 3} \mathbf{U}_{e 3} \quad (64)$$

Set a limit on the neutrino mixing angle $\sin \theta_{13}$ from the above muon decay bound.



Limits set on θ_{13} from the present and future bounds on the LFV decay $\mu \rightarrow e\gamma$. (a) $\tan\beta = 50$, $B = 5 \times 10^{-14}$, (b) $\tan\beta = 50$, $B = 1.2 \times 10^{-11}$, (c) $\tan\beta = 5$, $B = 1.2 \times 10^{-11}$, (d) $\tan\beta = 5$, $B = 5 \times 10^{-14}$, The area above(below) the curves is allowed(excluded).

Summary

- Casas-Ibara \mathbf{R} matrix is basis invariant and a general form of \mathbf{R} has been given.
- Two limiting $\text{SO}(10)$ models were presented and the interplay of $\sin \theta_{13}$ and $\mu \rightarrow e\gamma$ was studied.